



Max-Planck-Institut
für Radioastronomie

January 22nd, 2015
mm-VLBI workshop, Bologna
Eduardo Ros (MPIfR)

GMVA AND ALMA

The team



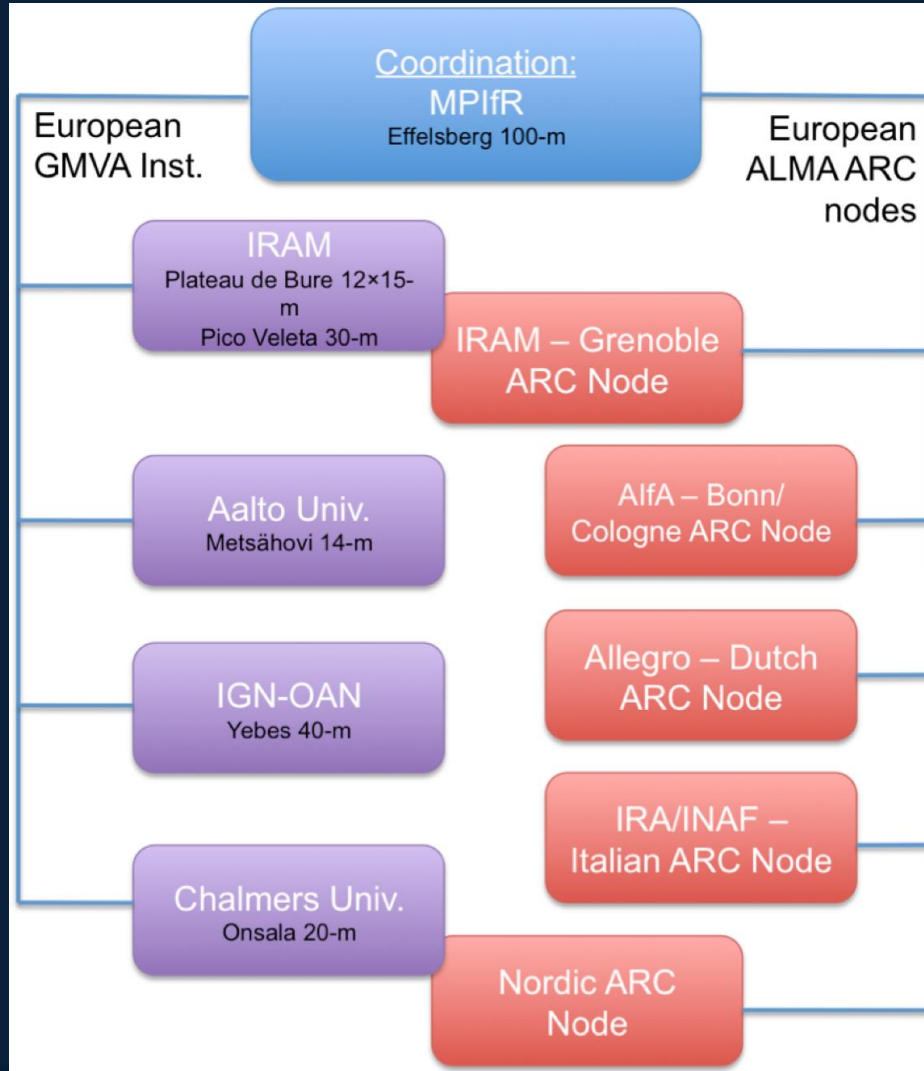
Laing
Testi

Bremer

Tornikoski

Bachiller
Colomer

Lindqvist
Martí-Vidal
Vlemmings



Alef
Bertarini
Krichbaum
Rottmann
Ros
Zensus

Mühle

Tilanus

Brand
Liuzzo
Massardi



The science

- Study of the immediate vicinity of the central engine in AGN
 - Schwarzschild-radius size imaging
 - Collimation and acceleration region in jets
- Spectral-line VLBI of absorbing systems for nature constant determination
- VLBI of masers in stellar objects and AGN
- Astrometry in the Milky Way and beyond



Background

1. ALMA upgraded by implementing a beamformer for phasing up ALMA for VLBI and PSR observations
2. The Event Horizon Telescope Experiment will perform 1-mm observations of SgrA* and M87
3. Preparing operation of ALMA in a VLBI network
 - GMVA operates at 3mm wavelength
4. Joint observations ALMA with VLBI networks



Why?

- ⦿ ALMA was de-scoped in the development phase
- ⦿ Hooks for phasing are in the correlator, ... for VLBI at a later date
- ⦿ That “later date” is NOW! Further additions:
 - H-Maser
 - Phasing system & Software
 - VLBI formatting, data transport & Recorders



Indeed, why?

- ⦿ 7mm VLBI: EVN+KVN tests performed
- ⦿ 3mm VLBI: GMVA, including KVN
 - Regular observations, though more difficult than at cm
 - Sensitivity sufficient for observing a few 100 sources
- ⦿ Tests at 2 and 1 mm, lack sensitivity – EHT observes 1x per year
 - Increase sensitivity with broader bandwidth (aim: 64 Gbps : 2x2x4 GHz) and with **phased arrays** !
(Add signals of antennas of a local array coherently)
 - Done at Plateau de Bure for mm-VLBI (GMVA)
 - Only for 2x128 MHz=256 MHz; still ~95% efficient at 1 mm
 - New NOEMA correlator on the horizon
 - CARMA, SMA, LMT, NOEMA...
- ⦿ ALMA !! (also offers very long **north-south** baselines)
 - EHT (1mm), GMVA (3mm), VLBA (7&3mm), EVN (7mm),...



Ongoing activities

- ALMA Phasing Program (phasing 2015/Q2)
- Science case for ALMA beamformer (Fish et al. 2013, arXiv:1309.3519)
- White paper on implementation (Tilanus et al. 2014, arXiv:1406.4650)
- ERC granted a Synergy Proposal (BlackHoleCam) to several European partners to achieve scientific goals
- **This proposal to define operations with ALMA**
- ALMA Cycle 4 call in Spring 2016 (GMVA calls in Aug15, Feb16, ...)



Layout

- ⦿ Network procedures
- ⦿ Proposal handling
- ⦿ Scheduling
- ⦿ Disk logistics, correlation, archiving
- ⦿ Data rights
- ⦿ User support
- ⦿ Resources & funding

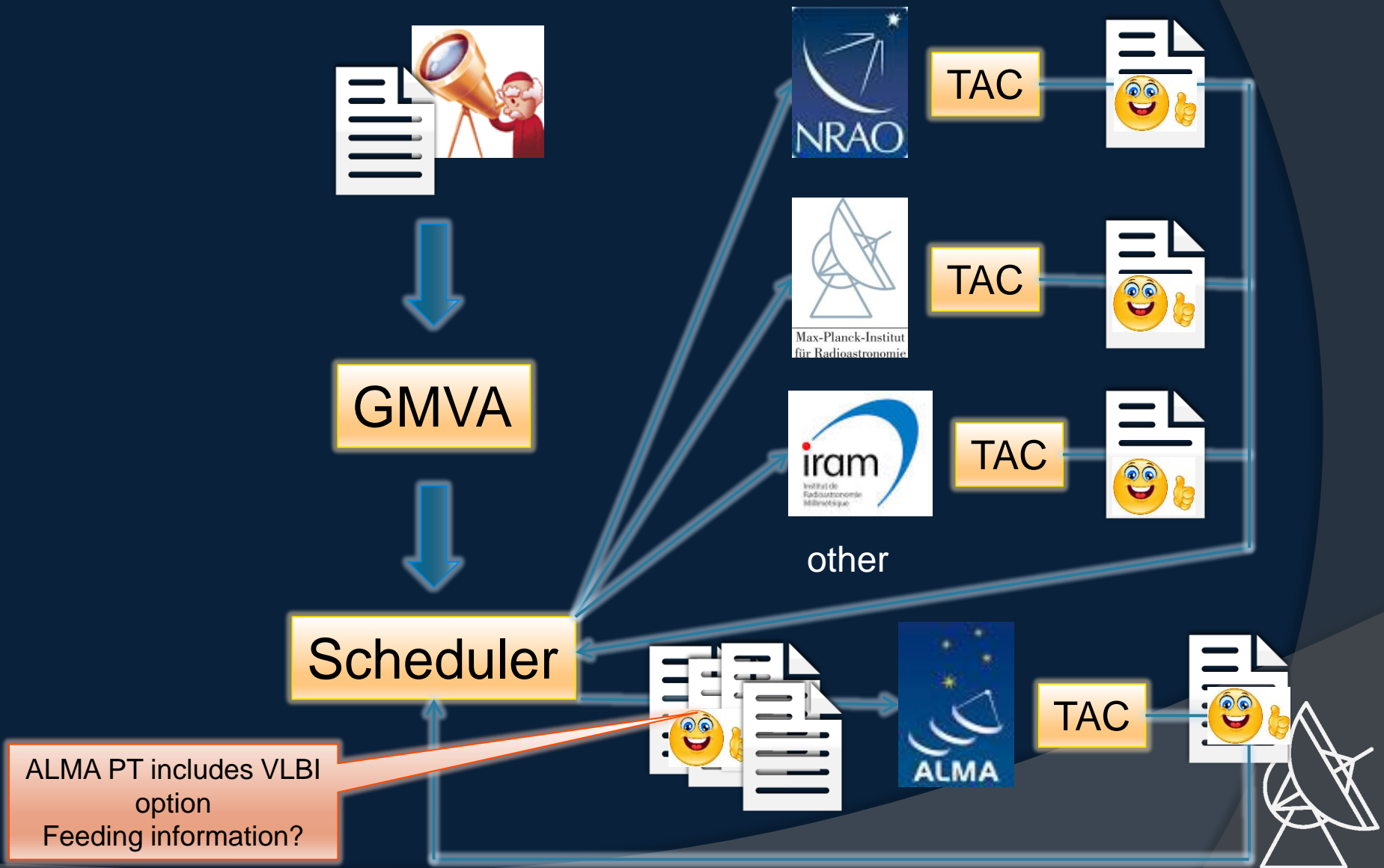


Network procedures

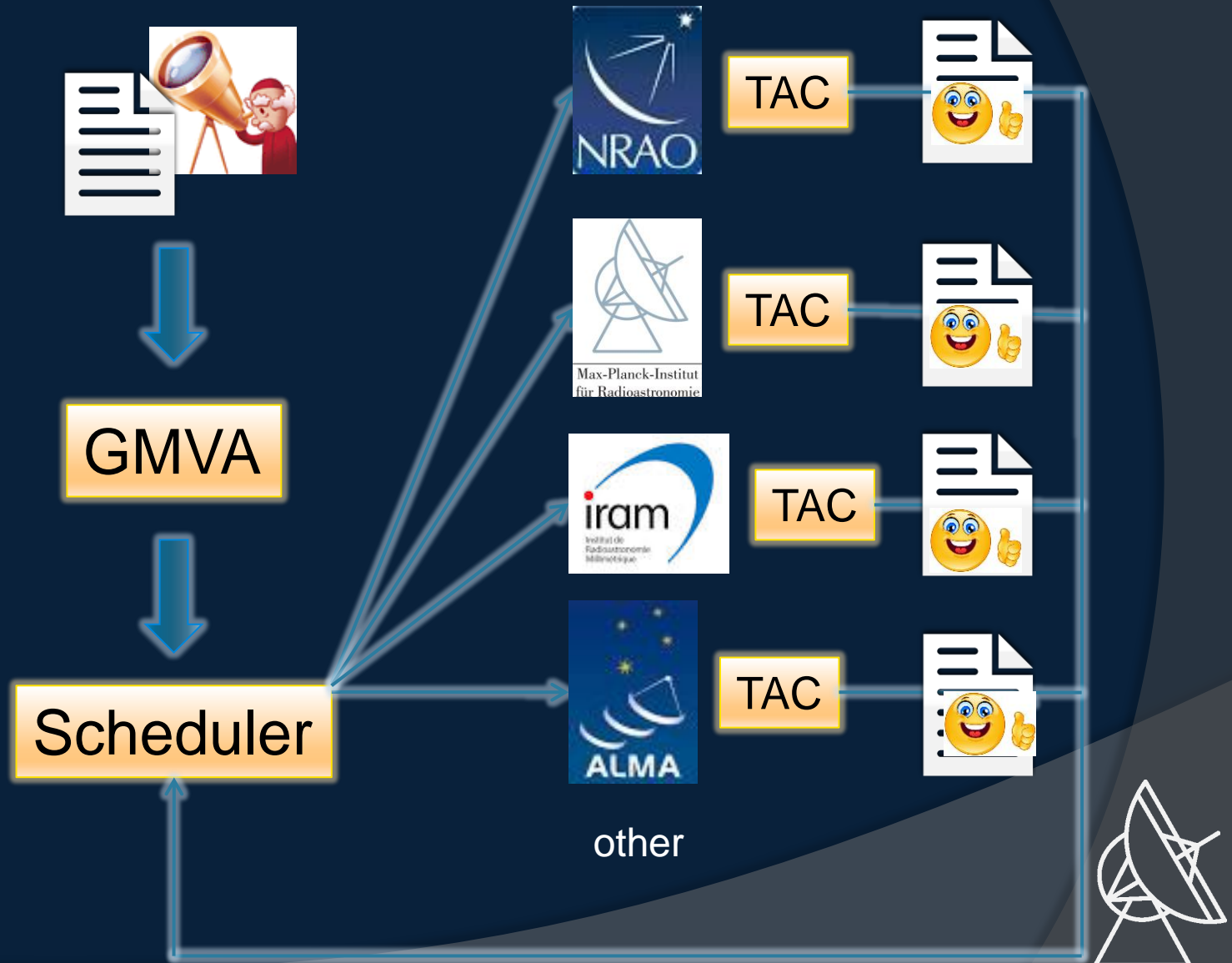
- ⦿ Formal agreement between participating sites: MoU
- ⦿ User-oriented infrastructure
- ⦿ Open skies, peer reviewed at each network site (or joint TAC if agreed)



Proposal handling



Proposal handling (alternative)



Scheduling

- ⦿ Considerations:
 - weather statistics for all sites for planning
 - weather decision (triggered observations)
 - telescope constrains (frequency agility)
- ⦿ ALMA observing:
 - manual scheduling possible
 - 0th-order approach: all stations use frequency and bit rate setup of ALMA
- ⦿ Suggestion:
 - agreement on a 'window of observation', with weather-triggered observation at a given GST range for several optional days



Disk logistics, correlation, archiving, and data rights

- Telescopes commit to ship the necessary media and to make available additional, if needed
- Small data portions sent per internet for fringe verification during observation
- Expert support available in the stations (friend of VLBI)
- Data will be correlated and stored at the MPIfR, with an archival copy at NRAO



Data rights

- ⦿ Present policy:
 - calibrator data are public
 - proprietary data: 1 yr after the observer gets the IDI-FITS file, after that, available at NRAO archive interface
- ⦿ Suggestion: similar policy for ALMA data and VLBI data including ALMA



Technical and user support

- ◉ GMVA: limited at present (open sky but expertise needed)
 - Schedule made for observers
 - Observing in correlation in absentia
 - Amplitude calibration checked for quality and consistency
 - Further support based on collaborations
- ◉ Potential ARC support at
 - Technical advice at proposal stage
 - ALMA calibration, availability of ALMA visibilities
 - Monitoring of VLBI calibrators
 - Support with data reduction (AIPS, CASA, etc.)



Funding

- Proposal handling
 - Data recording at telescopes
 - Correlation and data proof
 - Archiving and user support
 - VLBI-specific equipment maintenance
 - Media
-
- Typically covered locally ($\approx 10^4$ €/yr)
 - ALMA should fund it partially – Issue: friend of VLBI



Key issue:

Friend of VLBI at ALMA

Funding?

- ◉ Media shipping and organisation
- ◉ Conversion VEX into field system
 - schedules
 - check list
 - scan list
- ◉ Phasing/pointing
- ◉ Coordination with tech. staff/groups in telescope
- ◉ Post processing support: log files, calibration data, failure report
- ◉ Potential duties:
 - Station software, maintenance & upgrade VLBI hardware, network management, testing, of equipment/firmware/hardware
- ◉ Knows the institution from inside, it is not a “parachutist”
- ◉ Guarantees the success of VLBI observation by monitoring all the aspects involved



Summary

- Proposal of upgrading ALMA for VLBI observing procedures: exercise towards a global mm-VLBI array
- Getting all stakeholders and involving the whole ALMA community
- Immediate goal: proposal-based joint observations at the ALMA Cycle 4 (associated with the GMVA call for Feb16)
- Discussion of details on proposing/scheduling/data rights/user support involving the whole ALMA community

