

ALMA data reimaging

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1. Introduction

- Goal: create the complete set of PL Imaging products for all previous cycles
- Idea:
 - While redoing the calibration for all ES cycles seems impossible due to the very large amount of manual intervention required (redoing QA2)
 - Imaging, starting from the calibrated MS could actually work
 - Check the feasibility and cost

why

- It is clear that especially in interferometry it is not possible to create products that are suitable for all science cases (rebinning, tapering, change of resolution, weighting schemes, ...)
- In particular we believe that even if the ALMA archive had complete and science-grade data-products **still** a large fraction of users would want to reimage

- Why we nevertheless believe a full set of Images for ALMA is indispensable
 - Previews/Quick-look images can speed up the discovery by many orders of magnitude
 - We can measure fiducial rms and dynamic range from the images
 - The products can be sent to CARTA
 - We can do a generic post-analysis with ADMIT
 - We can offer the products to the VO
 - We can compute the HiPS and offer that in the AQ
 - If the products are too large for the users to create, we can create them for the users and provide cut-outs
 - The images can serve as starting point for small dedicated use reimaging
 - We can provide a homogeneous set of all ALMA images

- 1) Extract the data+scripts from the Archive
- 2) Discover the CASA version to use for the MS restore
- 3) run scriptForPI.py to get the calibrated MS
- 4) Apply known fixes to the data
- 5) Identify the MS(s) and run the ALMA Imaging PL **without changing any of the default parameters**
- 6) Run the Preview code
- 7) Run footprint extraction
- 8) Run ADMIT
- 9) Log the result of the run (errors, #products, ...) in a database

what this is *not*

- Evaluation of the scientific quality of the Imaging PL
- Creation of the best possible products
- Creation of multiple products for the different science cases

2. status

- Timeline
 - 2012-2013 First code to do PL calibration vs. manual calibration checking. Both were then imaged with the same parameters.
 - 2014-2016 Code to run the imaging PL on the restored MS
 - 2016-2017 Rewrote the code completely

- Code
 - Python
 - OO
 - Supports remote fetching of data through the RH
 - Supports ORACLE and sqlite3 databases
 - Works around most known issues (README files, file names, ...)
 - Cycle 0 as well as Cycle 1+
 - Runs with currently bleeding edge Imaging PL
 - Uses hifatargets.py
 - Makes Previews, footprint, runs ADMIT
 - Can use parallel CASA



hifatargets.py

```
def hifatargets (vislist, importonly=False, pipelinemode='automatic', interactive=True):  
...  
try:  
    h_init() # Initialize the pipeline  
    hifa_importdata (vis=vislist, dbservice=False, pipelinemode=pipelinemode) # Load the data  
    if importonly:  
        raise Exception(IMPORT_ONLY)  
    hif_mstransform (pipelinemode=pipelinemode) # Split out the target data  
    hifa_flagtargets (pipelinemode=pipelinemode) # Flag the target data  
    hif_makeimlist (specmode='mfs', pipelinemode=pipelinemode) # Make a list of mfs targets  
    hif_findcont(pipelinemode=pipelinemode) # Find continuum frequency  
    hif_uvcontfit(pipelinemode=pipelinemode) # Fit the continuum  
    hif_uvcontsub(pipelinemode=pipelinemode) # Subtract the continuum fit  
    hif_makeimages (pipelinemode=pipelinemode) # Make clean cont images  
    hif_makeimlist (specmode='cont', pipelinemode=pipelinemode) # Make a list of cont targets  
    hif_makeimages (pipelinemode=pipelinemode) # Make clean mfs images  
    hif_makeimlist (width="", pipelinemode=pipelinemode) # Make a list ofcontsub  
targets  
    hif_makeimages (subcontms=False, pipelinemode=pipelinemode) # Make clean cont subtracted  
    hif_exportdata(pipelinemode=pipelinemode) # ARI: also export the  
products
```

- From 2015-2017
 - 921 runs (with various versions of the PL)
 - 470 runs went completely through to ADMIT
- Typical errors
 - 75 memory error
 - 51 problems with the CASA versiof for restore
 - 37 no visibility found
 - 21 core dumps
 - ...

- weblog
- FITS files
- Footprints: 1 per FITS file
- Previews: 1 per FITS file
- ADMIT: 1 per 3D cube



products

casa_commands.log
casa_piperestorescript.py
casa_pipescript.py
cont.dat
flux.csv
oussid.NGC_1614_sci.spw0.cube.l.pb.fits.gz
oussid.NGC_1614_sci.spw0.cube.l.pbcor.fits.gz
oussid.NGC_1614_sci.spw0.mfs.l.pb.fits.gz
oussid.NGC_1614_sci.spw0.mfs.l.pbcor.fits.gz
oussid.NGC_1614_sci.spw0_1_2_3.cont.l.alpha.error.fits.gz
oussid.NGC_1614_sci.spw0_1_2_3.cont.l.alpha.fits.gz
oussid.NGC_1614_sci.spw0_1_2_3.cont.l.pb.tt0.fits.gz
oussid.NGC_1614_sci.spw0_1_2_3.cont.l.tt0.pbcor.fits.gz
oussid.NGC_1614_sci.spw0_1_2_3.cont.l.tt1.pbcor.fits.gz
oussid.NGC_1614_sci.spw1.cube.l.pb.fits.gz
oussid.NGC_1614_sci.spw1.cube.l.pbcor.fits.gz
oussid.NGC_1614_sci.spw1.mfs.l.pb.fits.gz
oussid.NGC_1614_sci.spw1.mfs.l.pbcor.fits.gz
oussid.NGC_1614_sci.spw2.cube.l.pb.fits.gz
oussid.NGC_1614_sci.spw2.cube.l.pbcor.fits.gz
oussid.NGC_1614_sci.spw2.mfs.l.pb.fits.gz
oussid.NGC_1614_sci.spw2.mfs.l.pbcor.fits.gz
oussid.NGC_1614_sci.spw3.cube.l.pb.fits.gz
oussid.NGC_1614_sci.spw3.cube.l.pbcor.fits.gz
oussid.NGC_1614_sci.spw3.mfs.l.pb.fits.gz
oussid.NGC_1614_sci.spw3.mfs.l.pbcor.fits.gz
pipeline_aquareport.xml
uid__A002_X8440e0_X4b44.ms.split.cal.calapply.txt
uid__A002_X8440e0_X4b44.ms.split.cal.flagversions.tgz
unknown.pipeline_manifest.xml
weblog.tgz

3. next steps

- Next steps
 - Build categories out of the holdings where we believe that the same results (success/failure & quality) will be reached
 - Run and verify e.g. 10 datasets of each category and validate/invalidate the approach
 - Look at the error-classes of the technical failures and see if they can be worked around.
 - Check for fixes that need to be potentially applied (e.g. CSV2555)

Summary

- A basis of runs as well as much of the code are in place
- Now comes the hard work