

# Archive data weblog and QA2 report

Obtaining information of the observation and  
calibration of ALMA Archive data



EUROPEAN ARC  
ALMA Regional Centre || Italian

# Purpose of ALMA weblog/QA2 report

The screenshot shows the ALMA web interface. At the top, there is a navigation bar with 'ALMA' logo, 'Home', 'By Topic', and 'By Task'. Below this, there are two main sections: 'Observation Overview' and 'Pipeline Summary'. The 'Observation Overview' section contains a table with the following data:

Project	uid://A001/X10e/X13a
Principal Investigator	jeskj
Observation Start	2015-05-17 06:09:08 UTC
Observation End	2015-05-17 06:46:20 UTC

The 'Pipeline Summary' section contains a table with the following data:

Pipeline Version	31667 (Pipeline-Cycle2-R1-B)
CASA Version	4.2.2 r30986
Pipeline Start	2015-07-15 16:32:03 UTC
Execution Duration	5:36:43

Below these sections is the 'Observation Summary' section, which contains a table with the following data:

Measurement Set	Receivers	Num Antennas	Time (UTC)			Baseline Length			Size
			Start	End	On Source	Min	Max	RMS	
Observing Unit Set Status: unknown Scheduling Block ID: uid://A001/X120/Xba									
Session: session_2									
uid__A002_Xa0b40d_X3cb8.ms	ALMA Band 7	36	2015-05-17 06:09:08	2015-05-17 06:46:19	0:13:43	21.4 m	555.5 m	221.7 m	23.2 GB

- Information about the observation: weather, antenna configuration, observation setup and strategy
- Information about the calibration of the data: contains plots of various calibration steps that allow you to check the calibration
- When pipeline reduced >> weblog. When reduced manually (mostly Cycle I and older data), you find the QA2 report (plots and txt file).
- In future: more data reduced by pipeline, weblog will become common for most datasets

# Location of the weblog/QA2 report?

If you download the observation product the weblog will be in

`project_code/  
science_goal.sous/  
group.gous/  
member.mous/  
qa/`

weblog :  
gunzip and untar weblog.tar.gz  
cd weblog-date/html/  
and open in browser index.html

QA2 report:

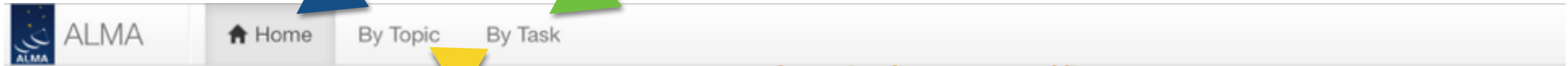
`uid*_qa2_part1.png, uid*_qa2_part2.png, uid*_qa2_part3.png, uid*_textfile.txt`  
(uid\* will be the execution block names)

In our example:

```
project_code=2013.1.00278.S  
sous=uid___A00I_X120_X100  
gous=uid___A00I_X120_X101  
mous=uid___A00I_X120_X102  
date=20150715T163202
```

# Weblog home

calibration task and their products



summary of tasks/warning/flags

## Observation Overview

Project	uid://A001/X10e/X13a
Principal Investigator	jeskj
Observation Start	2015-05-17 06:09:08 UTC
Observation End	2015-05-17 06:46:20 UTC

## Pipeline Summary

Pipeline Version	31667 (Pipeline-Cycle2-R1-B)
CASA Version	4.2.2 r30986
Pipeline Start	2015-07-15 16:32:03 UTC
Execution Duration	5:36:43

scriptForPI.py needs to be run in this Casa version

## Observation Summary

for the list of execution blocks (in this case there is only one EB)

Measurement Set	Receivers	Num Antennas	Time (UTC)			Baseline Length			Size
			Start	End	On Source	Min	Max	RMS	
Observing Unit Set Status: unknown Scheduling Block ID: uid://A001/X120/Xba									
Session: session_2									
<a href="#">uid__A002_Xa0b40d_X3cb8.ms</a>	ALMA Band 7	36	2015-05-17 06:09:08	2015-05-17 06:46:19	0:13:43	21.4 m	555.5 m	221.7 m	23.2 GB

Click for more info!

baseline lengths resolution

SESSION 'SESSION\_2'

**uid\_\_A002\_Xa0b40d\_X3cb8.ms**

# Overview of 'uid\_\_A002\_Xa0b40d\_X3cb8.ms'

(the execution block)

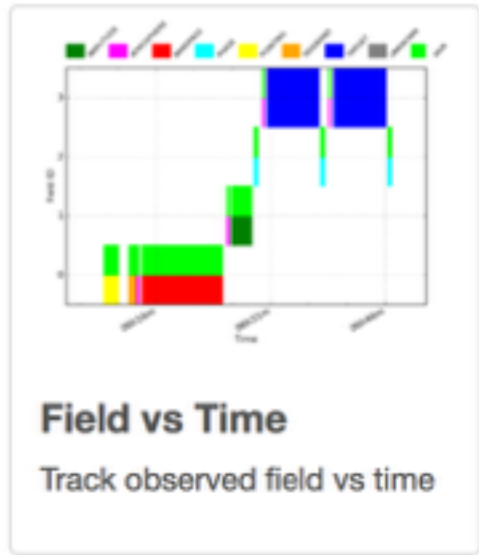
Blue titles can be clicked for having more info

## Observation Execution Time

Start Time	2015-05-17 06:09:08
End Time	2015-05-17 06:46:19
Total Time on Source	0:32:29
Total Time on Science Target	0:13:43

listobs output

Observation summary file: incl. obs. schedule, source coordinates, spectral setup, antenna positions...



## Spatial Setup

Science Targets	'IRAS16293-2422'
Calibrators	'J1517-2422', 'J1625-2527' and 'Titan'

## Spectral Setup

All Bands	'ALMA Band 7' and 'WVR'
Science Bands	'ALMA Band 7'

## Antenna Setup

Min Baseline	21.4 m
Max Baseline	555.5 m

## Sky Setup

Min Elevation	N/A
Max Elevation	N/A

## Weather



## Scans

Observation field of view is  $1.13 \times \lambda/D$  [rad]

for  $D=12m$ ,  $\nu=335.5$  GHz,  $\lambda=0.0894$  cm

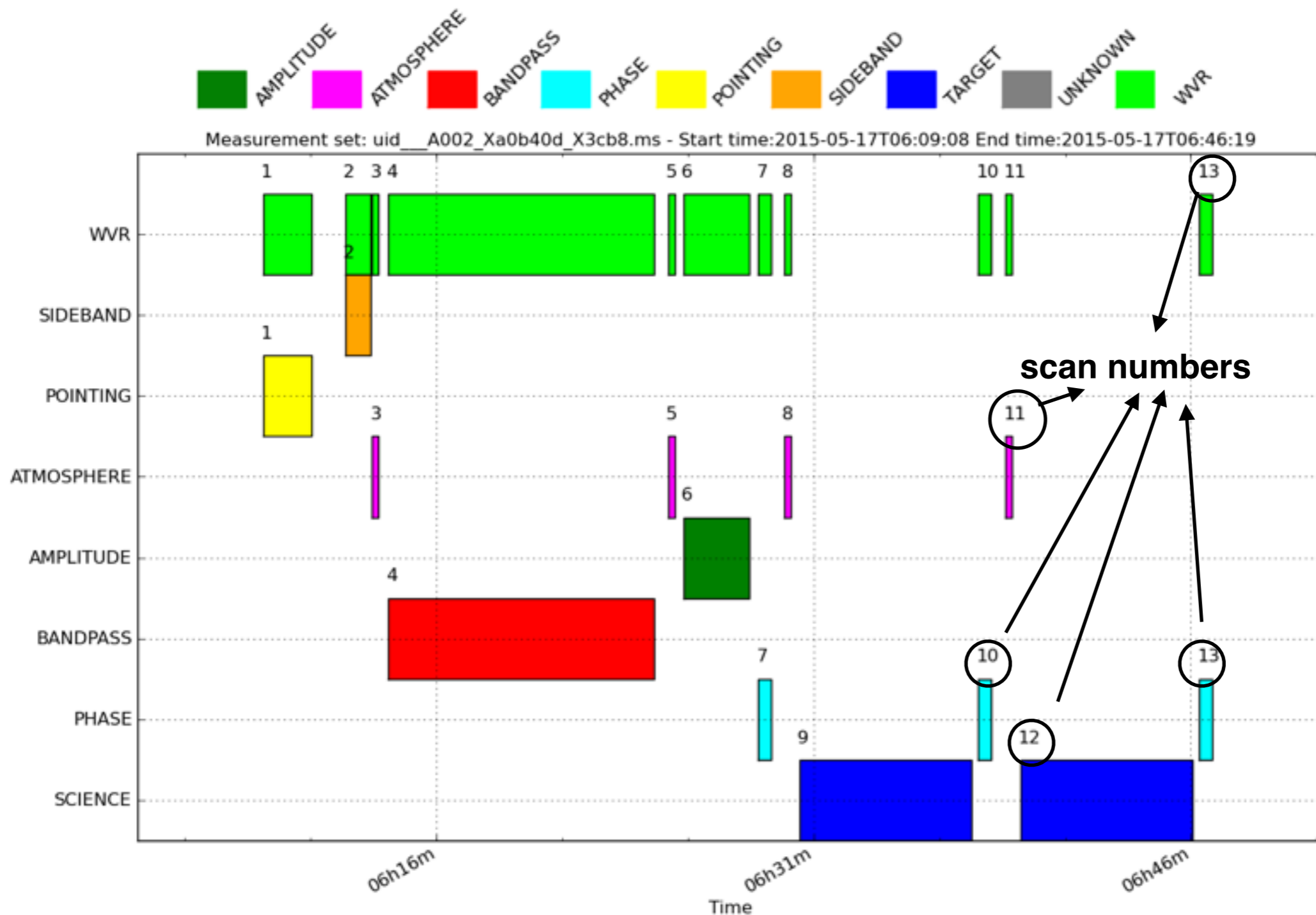
FOV=17.4"



# Observation intent vs. time

Shows the schedule of the observing track

WVR is also done during science observations!

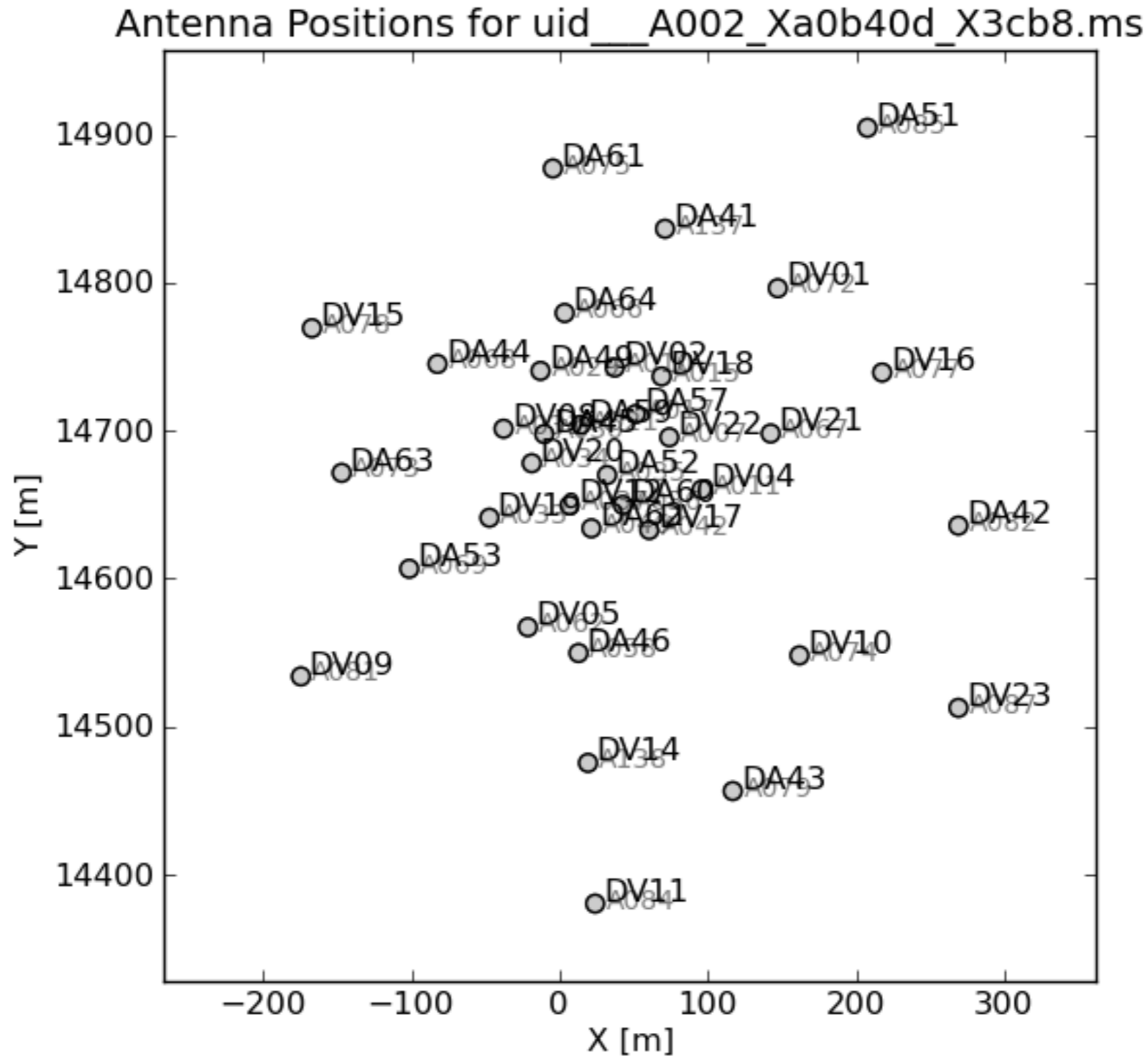






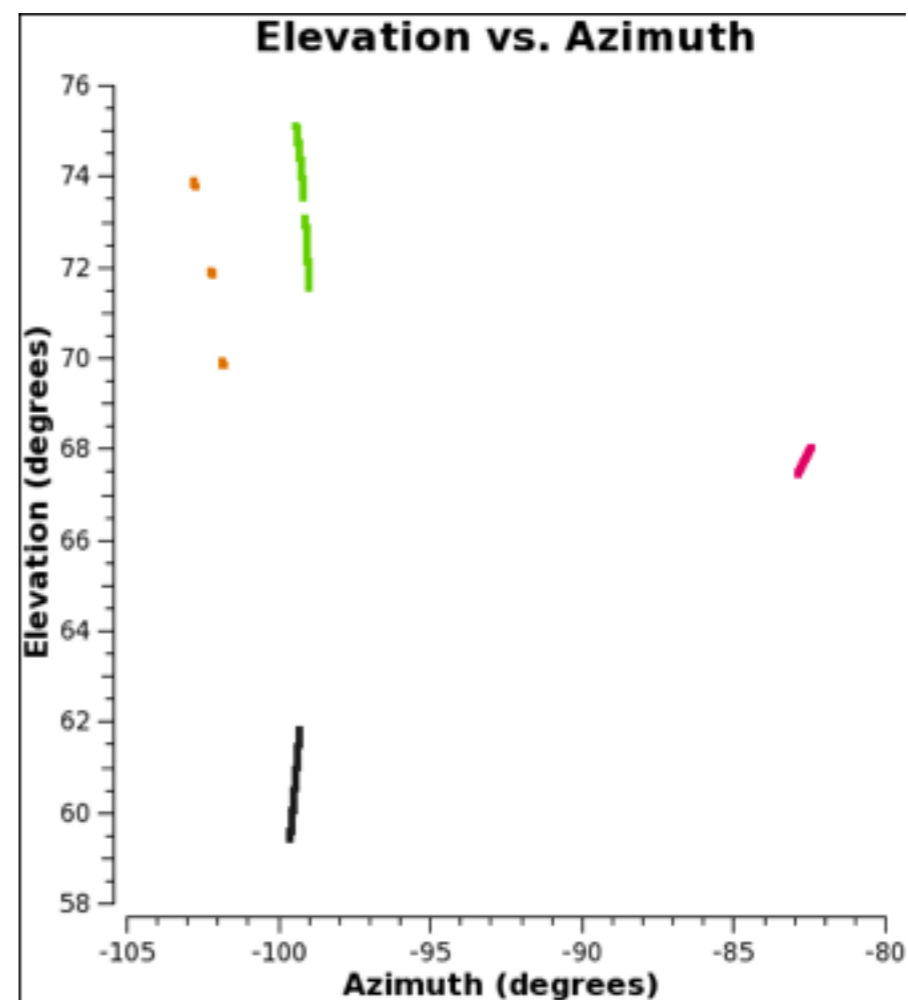
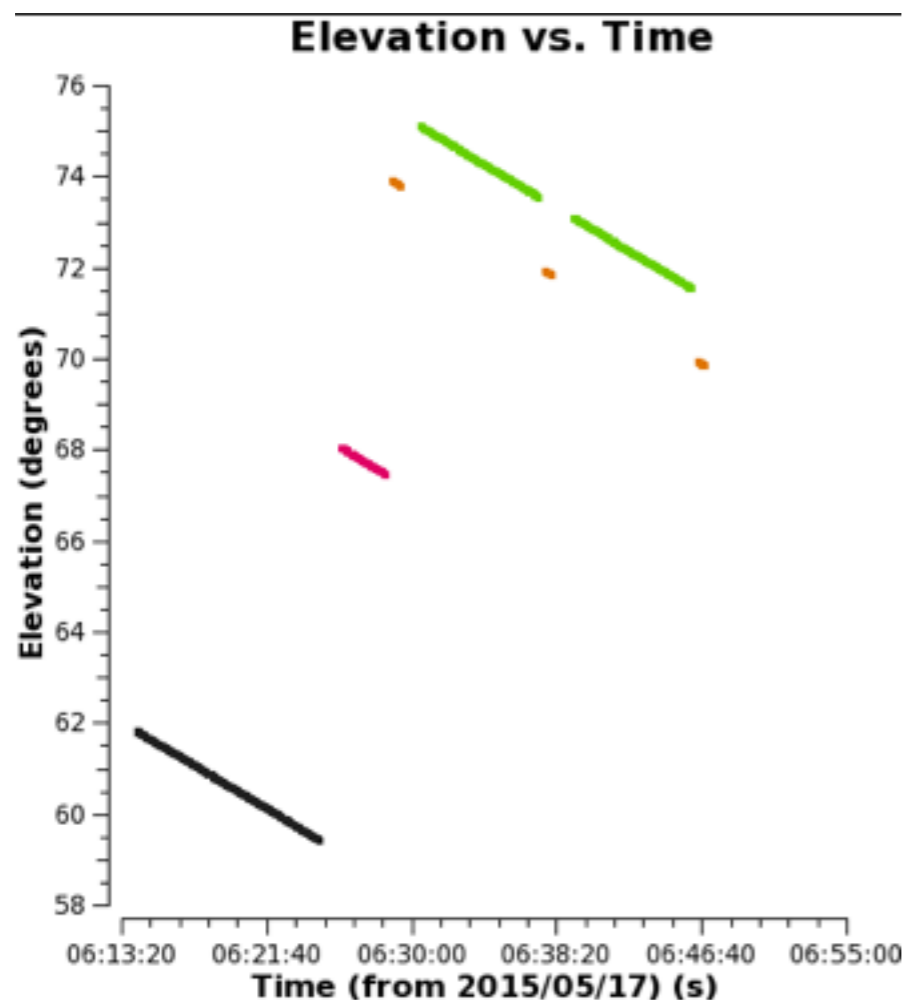
# Antenna setup

Antenna configuration plot, see also the baselines for determining the angular resolution and largest angular scale



# Sky setup plots

- In our case not present in **Sky Setup** of weblog, we can plot them ourselves in **plotms**:







### Warnings and Errors

Stage	Task	Type	Message
6	<a href="#">hifa_tsysflag</a>	Warning	flag edgechans - uid__A002_Xa0b40d_X3cb8.ms iteration 1 raised 12 flagging commands
6	<a href="#">hifa_tsysflag</a>	Warning	flag birdies - uid__A002_Xa0b40d_X3cb8.ms iteration 1 raised 2 flagging commands
11	<a href="#">hif_bpflagchans</a>	Warning	uid__A002_Xa0b40d_X3cb8.ms iteration 1 raised 1 flagging commands

### Tasks by Topic

Topic	Lowest Scoring Task	Min Score
Data Sets	1. <a href="#">hif_importdata</a> : Register measurement sets with the pipeline	1.00
Calibration	13. <a href="#">hifa_timegaincal</a> : Gain calibration	X-Y deviation  1.00
Flagging	14. <a href="#">hif_applycal</a> : Apply calibrations from context	22.36% data flagged  0.81
Imaging	No scoring tasks in this topic	N/A
Miscellaneous	4. <a href="#">hif_refant</a> : Select reference antennas	1.00

### Flagging Summaries

summaries are per source:  
 100= completely flagged,  
 0= no flags

uid\_\_A002\_Xa0b40d\_X3cb8.ms

spw	DA41	DA42	DA43	DA44	DA45	DA46	DA49	DA51	DA52	DA53	DA57	DA59	DA60	DA61	DA62	DA63	DA64	DV01	DV02
17	10.22	10.22	10.22	10.22	100.00	10.22	10.22	10.22	10.22	10.22	10.22	10.22	10.22	10.22	10.22	10.22	10.22	13.33	10.22
19	10.25	10.25	10.25	10.25	100.00	10.97	10.25	10.25	10.25	10.25	10.25	10.25	10.25	10.25	10.25	10.25	10.25	13.36	10.25
21	100.00	15.67	100.00	15.67	100.00	15.67	15.67	15.67	15.67	15.67	15.67	15.67	15.67	15.67	15.67	15.67	15.67	18.67	15.67
23	10.22	10.22	10.22	10.22	100.00	10.22	10.22	10.22	10.22	10.22	10.22	10.22	10.22	10.22	10.22	10.22	10.22	13.33	10.22

Flagging percentages for Source name: Titan, Intents: WVR,ATMOSPHERE,AMPLITUDE

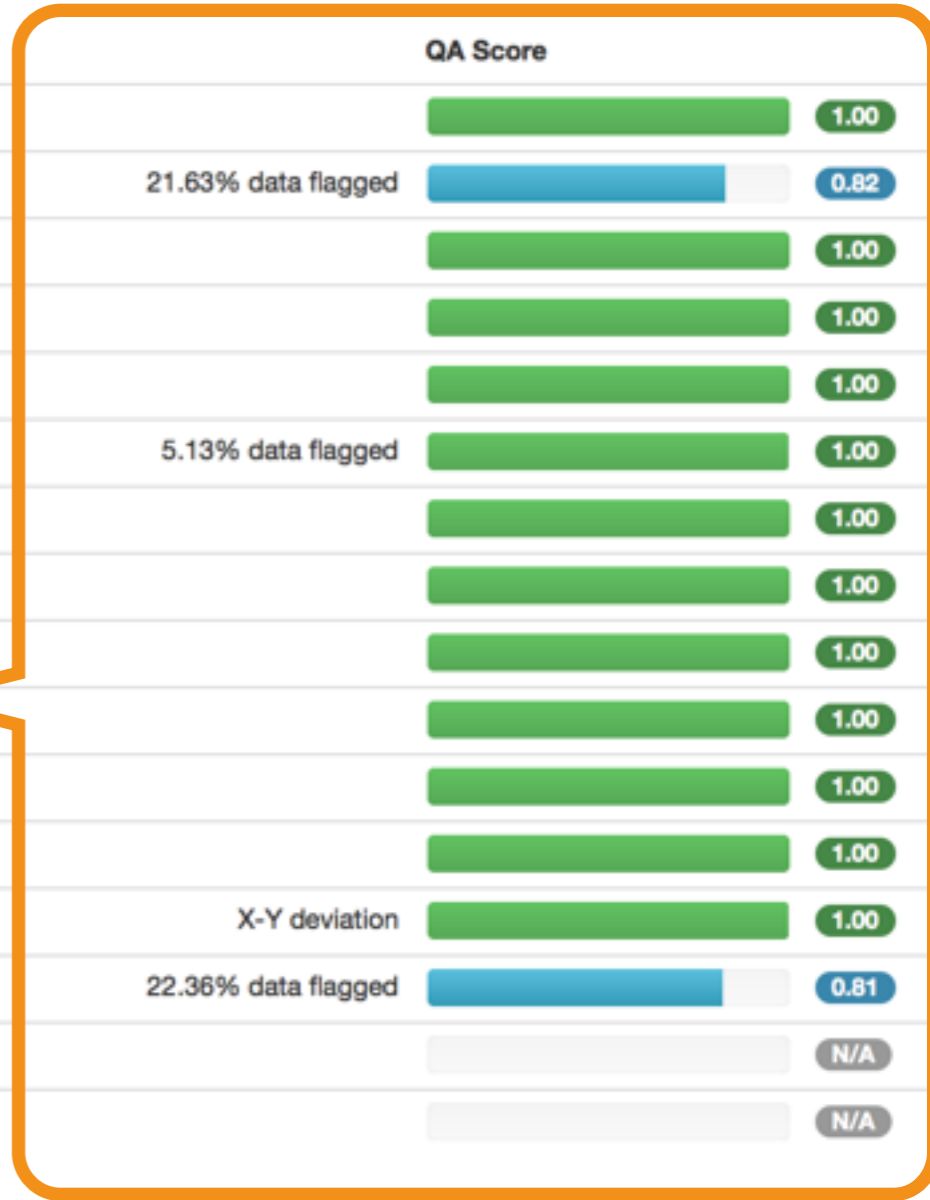
spw	DA41	DA42	DA43	DA44	DA45	DA46	DA49	DA51	DA52	DA53	DA57	DA59	DA60	DA61	DA62	DA63	DA64	DV01	DV02
17	8.53	10.31	8.53	8.53	100.00	10.31	8.53	12.17	10.28	10.36	9.42	11.25	8.53	8.53	9.42	8.53	10.25	12.58	12.56
19	8.55	10.33	8.55	8.55	100.00	11.05	8.55	12.19	10.30	10.38	9.44	11.27	8.55	8.55	9.44	8.55	10.27	12.60	12.58
21	13.89	15.56	100.00	13.89	100.00	15.56	13.89	17.31	15.53	15.61	14.72	16.44	13.89	13.89	14.72	13.89	15.50	17.72	17.64

Score	Colour	Comment
0.90-1.00	Green	Standard/Good
0.66-0.90	Blue	Below standard
0.33-0.66	Yellow	Warning
0.00-0.33	Red	Error

# Task Summary

- Task**
1. [hif\\_importdata](#): Register measurement sets with the pipeline
  2. [hifa\\_flagdata](#): ALMA deterministic flagging
  3. [hifa\\_fluxcalflag](#): Flag spectral features in solar system flux calibrators
  4. [hif\\_refant](#): Select reference antennas
  5. [hifa\\_tsyscal](#): Calculate Tsys calibration
  6. [hifa\\_tsysflag](#): Flag Tsys calibration
  7. [hifa\\_wvrgcalflag](#): Calculate and flag WVR calibration
  8. [hif\\_lowgainflag](#): Flag antennas with low gain
  9. [hif\\_setjy](#): Set calibrator model visibilities
  10. [hif\\_bandpass](#): Bandpass calibration
  11. [hif\\_bpflagchans](#): Flag channels of bandpass calibration
  12. [hifa\\_gfluxscale](#): Transfer fluxscale from amplitude calibrator
  13. [hifa\\_timegaincal](#): Gain calibration
  14. [hif\\_applycal](#): Apply calibrations from context
  15. [hif\\_makecleanlist](#): Compile a list of cleaned images to be calculated
  16. [hif\\_cleanlist](#): Calculate clean products

steps in calibration pipeline



Quality assessment score

- CASA logs and scripts**
- download [casapy-20150715-163158.log](#) (10.5 MB) - logger output from pipeline execution
  - download [casa\\_commands.log](#) (788.2 KB) - casa commands used in pipeline
  - download [casa\\_pipescript.py](#) (2.0 KB) - script with pipeline commands

Pipeline scripts/commands, can be adapted (see doc)

Documentation: ALMA Science Pipeline QuickStart Guide and Reference Manual on <https://almascience.eso.org/documents-and-tools>

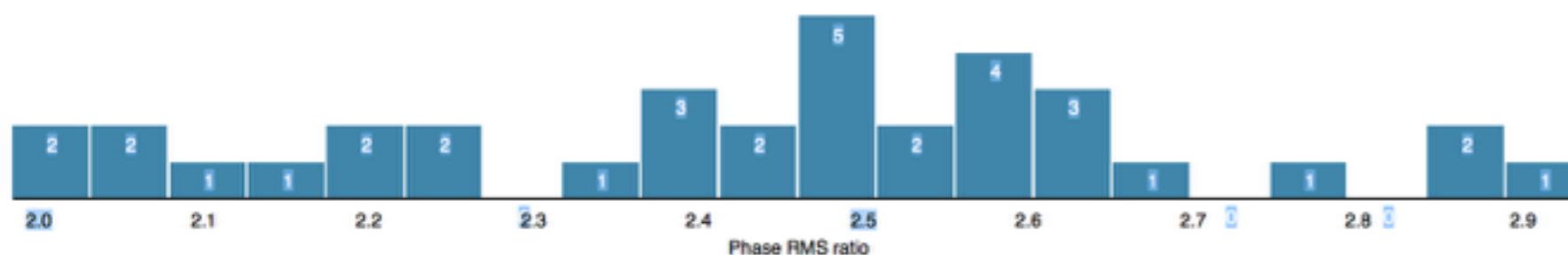
# WVR phase correction

Step `hifa_wvrgcalflag`: shown phases before and after WVR correction for a spw for all antennas:

One can go through all plots per antenna

WVR correction: phases flatter and closer to zero

RMS before/after histogram





# Tsys plots per spw

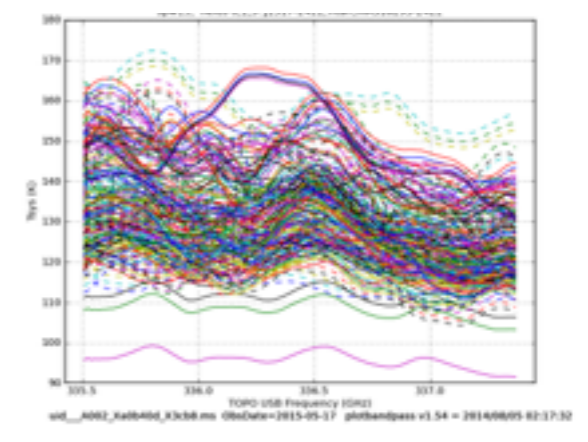
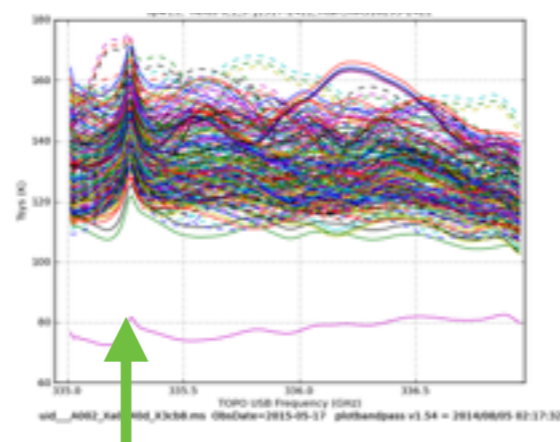
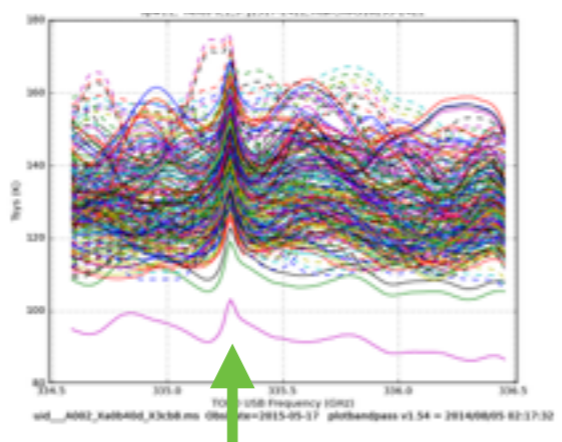
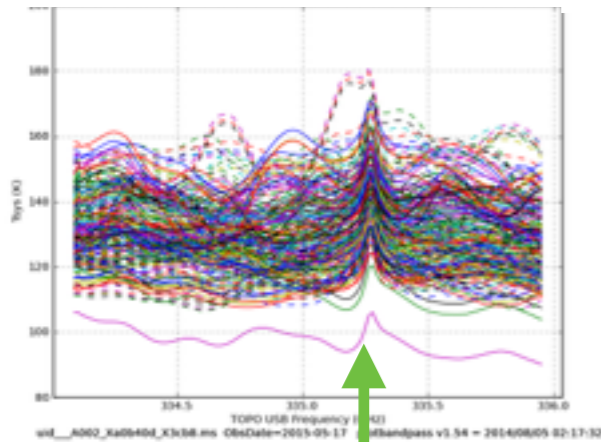
Step hifa\_tsysflag

SPW 9

SPW 11

SPW 13

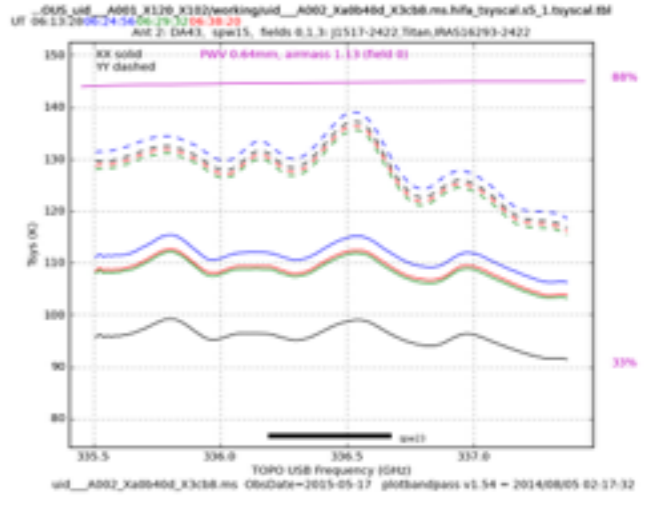
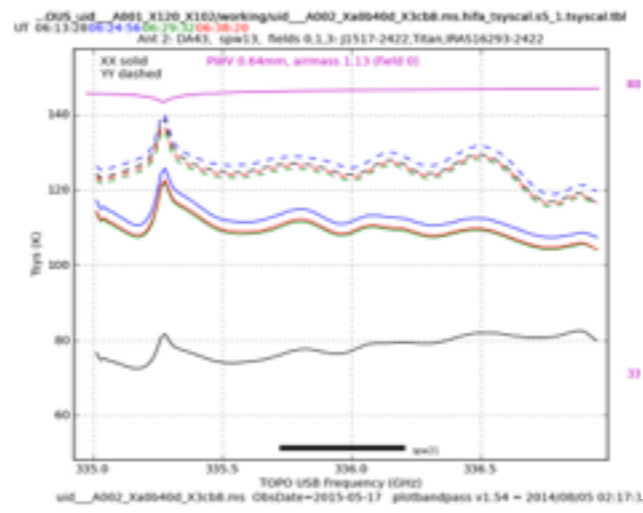
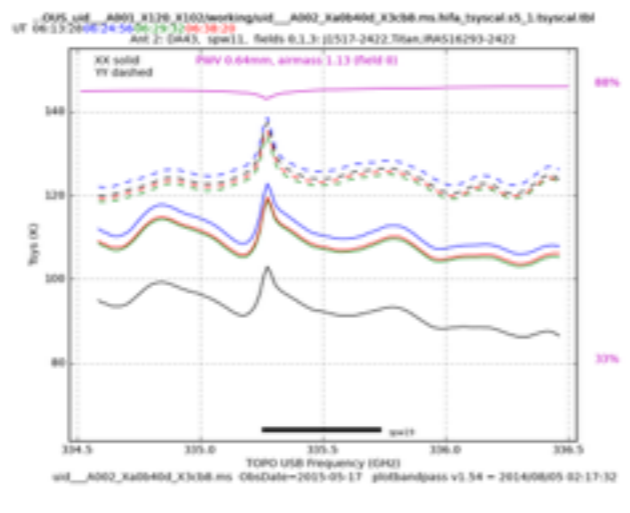
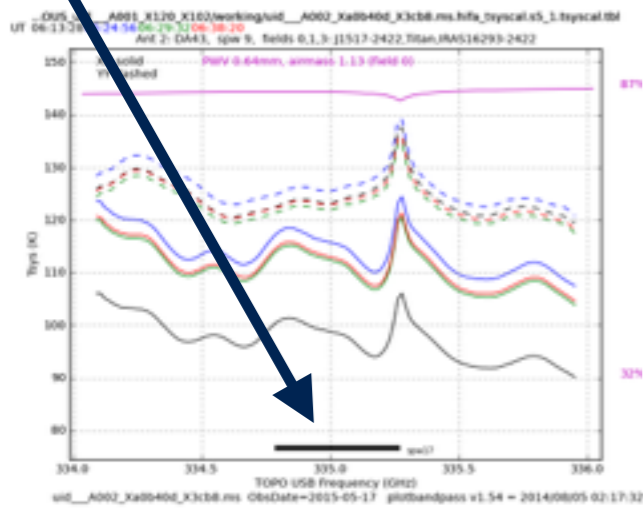
SPW 15



Atmospheric lines, of which most are known - use showatm=True in plotbandpass task to overlay atmospheric transmission curve or check single antenna Tsys plots

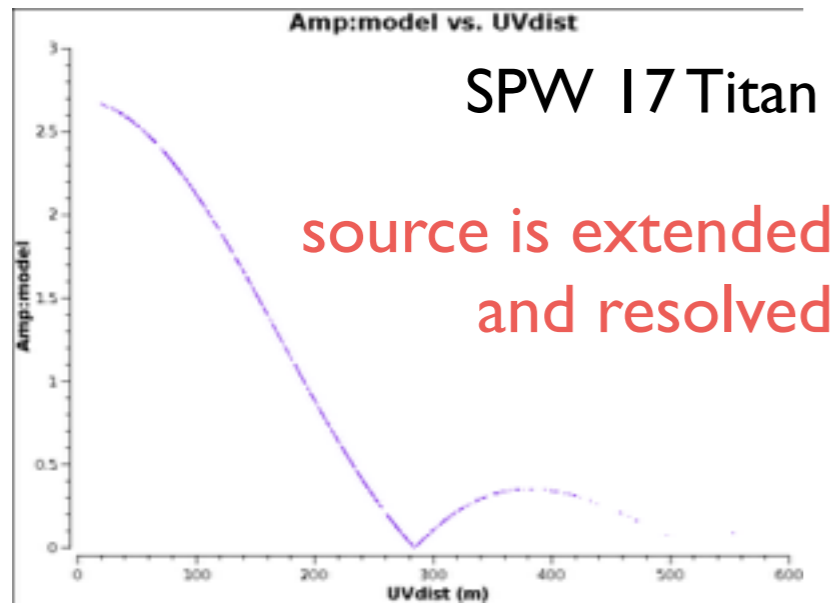
SPW 9, 11, 13, 15 for Ant DA41

black bar = science spectral window

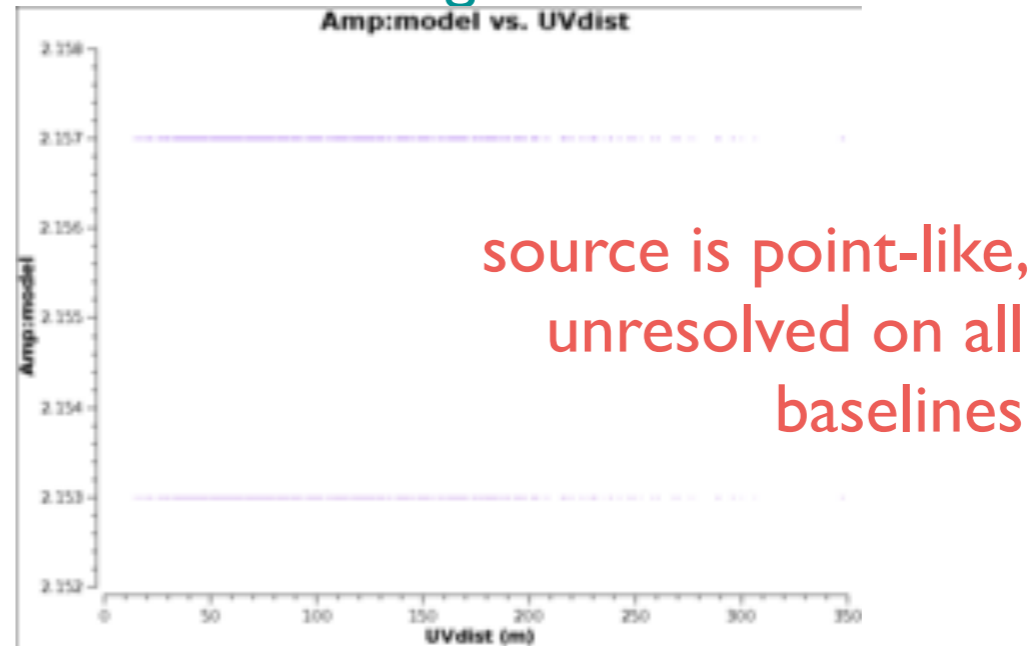


# UVrange flux calibrator

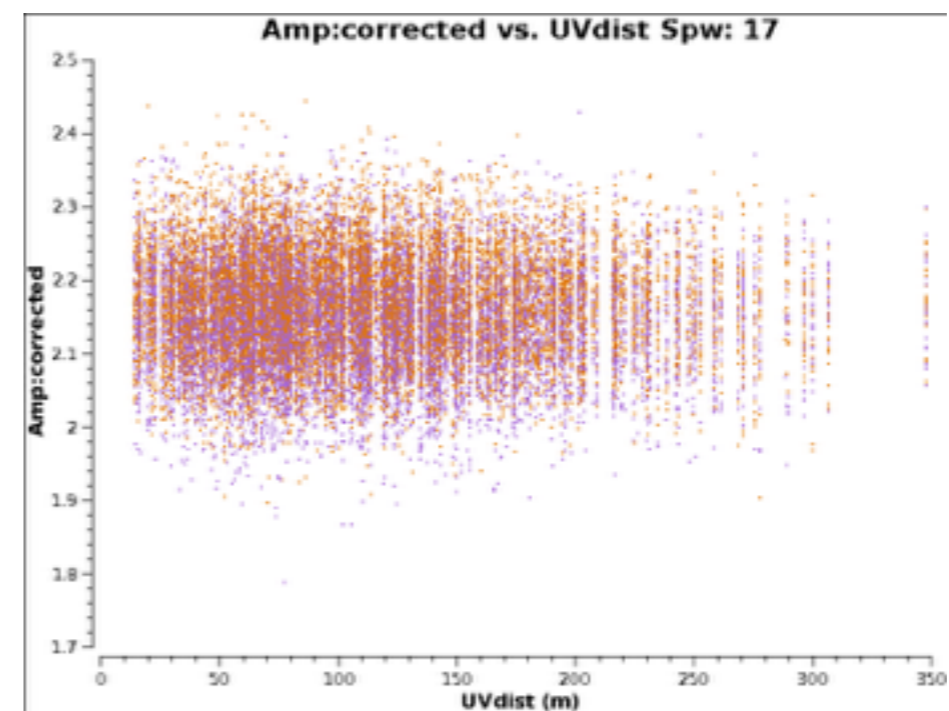
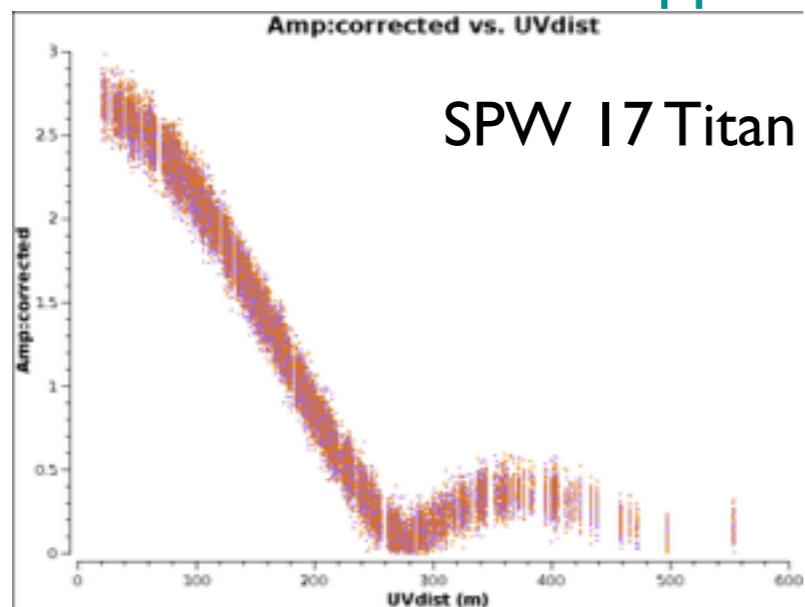
Step `hif_setjy`: For a Solar System flux calibrator a source model is used to obtain the flux measure



Step `hif_setjy`: For a Quasar flux calibrator, the flux measure is taken from last monitoring session



Step `hif_applycal`: uvdata flux calibrator with model applied



# Computed flux densities

- TASKS IN EXECUTION ORDER
1. hif\_importdata
  2. hifa\_flagdata
  3. hifa\_fluxcalflag
  4. hif\_refant
  5. hifa\_tsyscal
  6. hifa\_tsysflag
  7. hifa\_wvrgcalflag
  8. hif\_lowgainflag
  9. hif\_setjy
  10. hif\_bandpass
  11. hif\_bpflagchans
  - 12. hifa\_gfluxscale**
  13. hifa\_timegaincal
  14. hif\_applycal
  15. hif\_makecleanlist
  16. hif\_cleanlist

## 12. Phased-up fluxscale

Because Titan's (flux cal) emission was extended only the inner antennas are used for fluxscale determination

Back

### Results

#### Antennas Used for Flux Scaling

The following antennas were used for flux scaling, entries for unresolved flux calibrators are blank

Measurement Set	Antennas
uid___A002_Xa0b40d_X3cb8.ms	DV19, DV18, DA64, DA49, DA62, DA60, DA45, DA44, DA46, DA41, DV14, DV17, DV16, DA63, DV21, DV10, DA52, DA53, DV22, DV12, DA57, DV08, DV09, DV04, DA59, DV02, DV01, DV20, DV05

Antennas for Flux Calibration

#### Computed Flux Densities

The following flux densities were set in the measurement set model column and recorded in the pipeline context:

Measurement Set	Field	SpW	Flux Density			
			I	Q	U	V
uid___A002_Xa0b40d_X3cb8.ms	J1517-2422 (#0)	17	1.142 Jy ± 3.968 mJy (0.3%)	0.000 Jy	0.000 Jy	0.000 Jy
		19	1.149 Jy ± 5.600 mJy (0.5%)			
		21	1.145 Jy ± 13.350 mJy (1.2%)			
		23	1.136 Jy ± 4.577 mJy (0.4%)			
	J1625-2527 (#2)	17	686.225 mJy ± 3.385 mJy (0.5%)			
		19	685.805 mJy ± 4.784 mJy (0.7%)			
		21	684.340 mJy ± 10.235 mJy (1.5%)			
		23	678.322 mJy ± 3.097 mJy (0.5%)			

Phased-up Fluxscale Results

Resulting flux densities table of the bandpass and phase calibrators (QSOs) per spw.

Compare with ALMA CSC fluxes

## ALMA Calibrator Source Catalogue

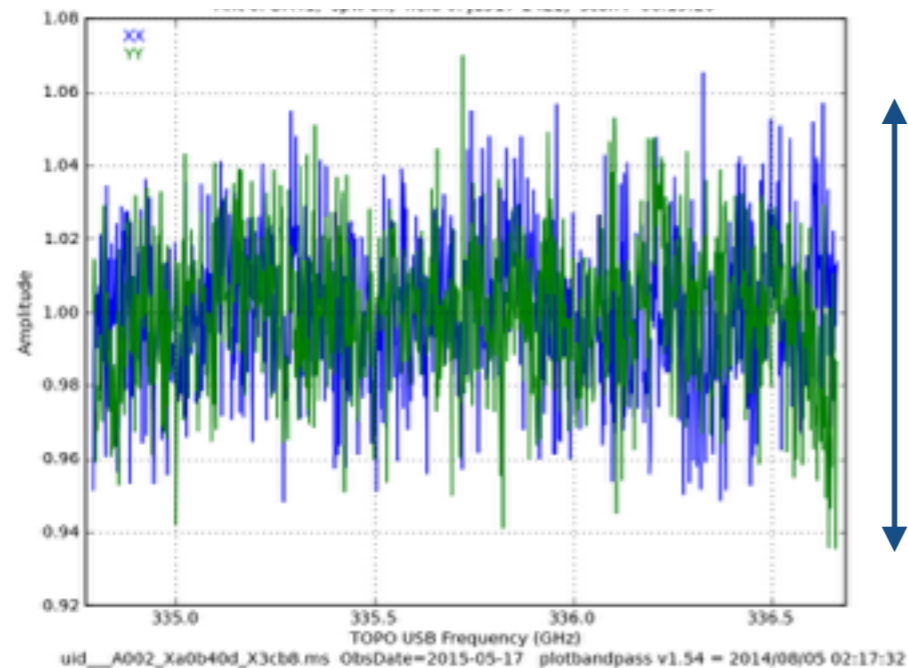
Name	RA	RA Err.	DEC	DEC Err.	Freq.	Band	Flux	Flux Err.	UvMin	UvMax	Observed
J1625-2527 J162...	16:25:46.8916	±0.0001	-25:27:38.327	±0.0001	285.0	7	0.658	±0.07		> 300.0	2014-09-18



# Bandpass: amp & phase vs freq plots

Steps `hifa_bandpass` `hif_applycal`

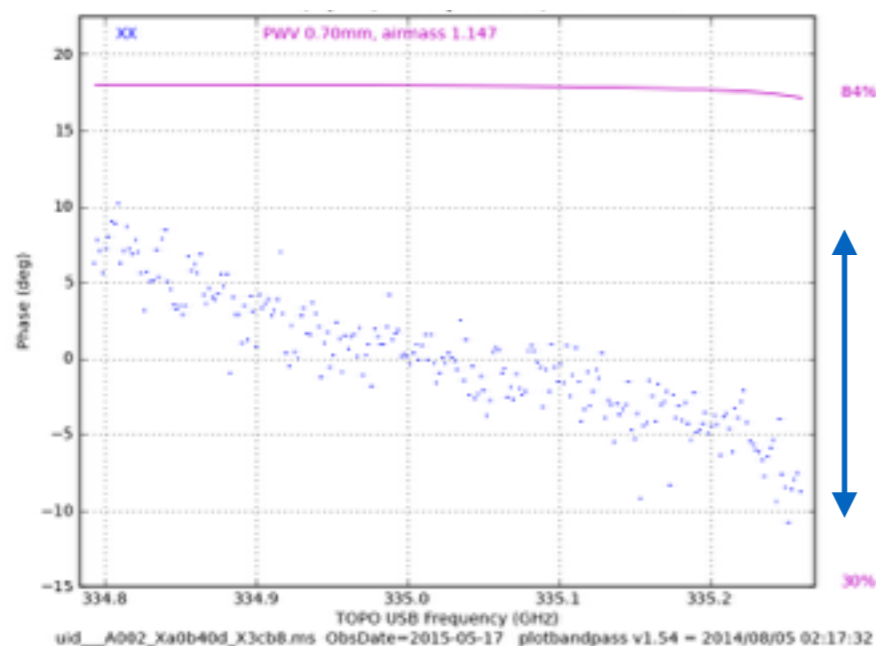
Amp vs Freq: ant DA4I, all spw



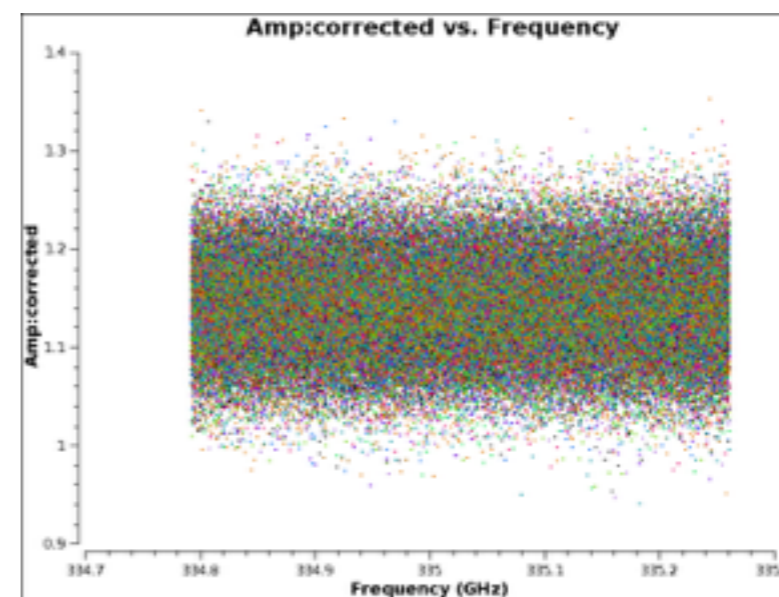
For each antenna one can look at the bandpass solutions, where a smooth fit and few noise are the way to recognize good quality solutions.

In Step `hif_applycal` one can see the ampl calibration of bandpass calibrator (all antennas shown) per spw. More elaborate plots can be done with `plotms`.

Phase vs Freq: ant DA4I, spw 17



`hif_applycal`: SPW 17, all antennas

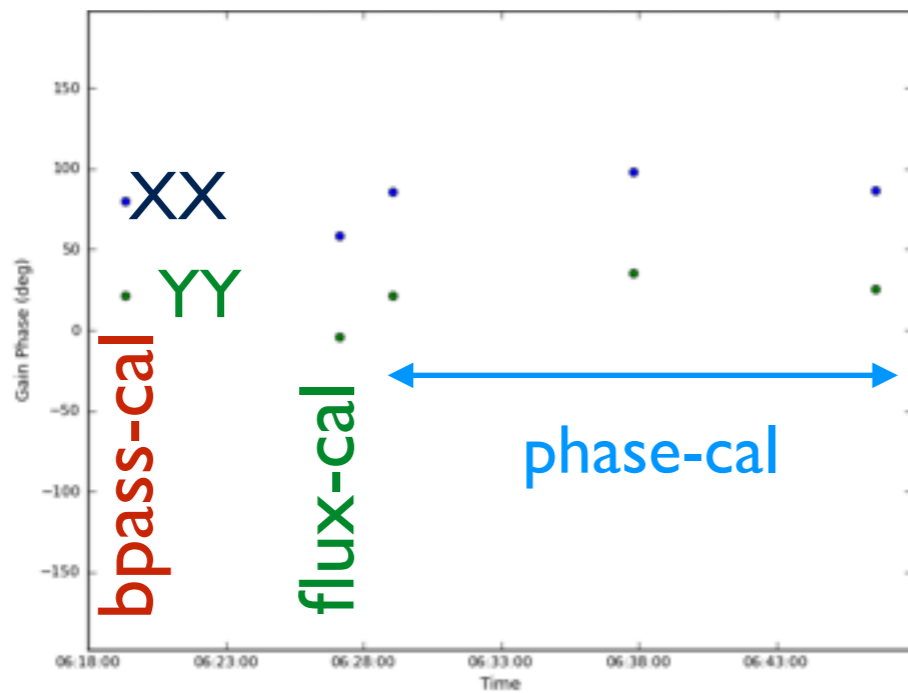




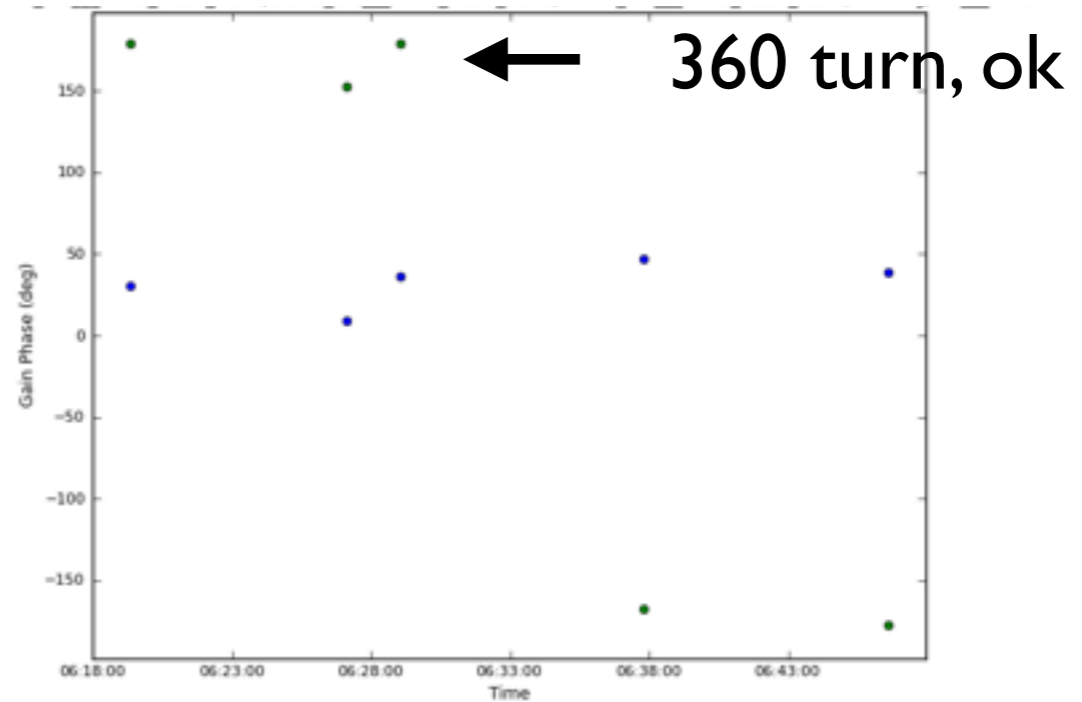
# Calibration tables: phase vs time

Step `hifa_timegaincal`

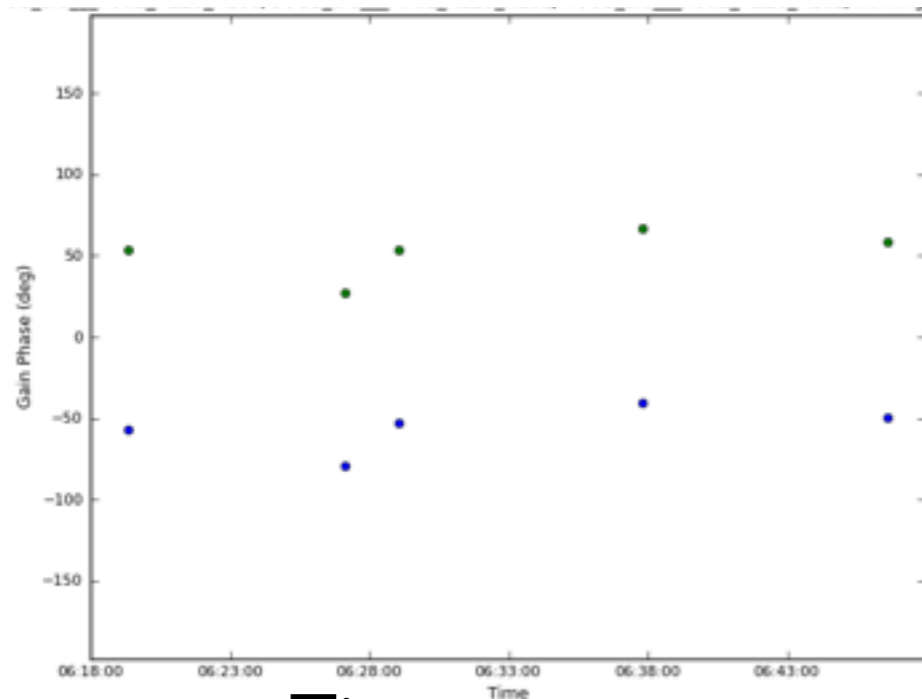
SPW 17,Ant DV19



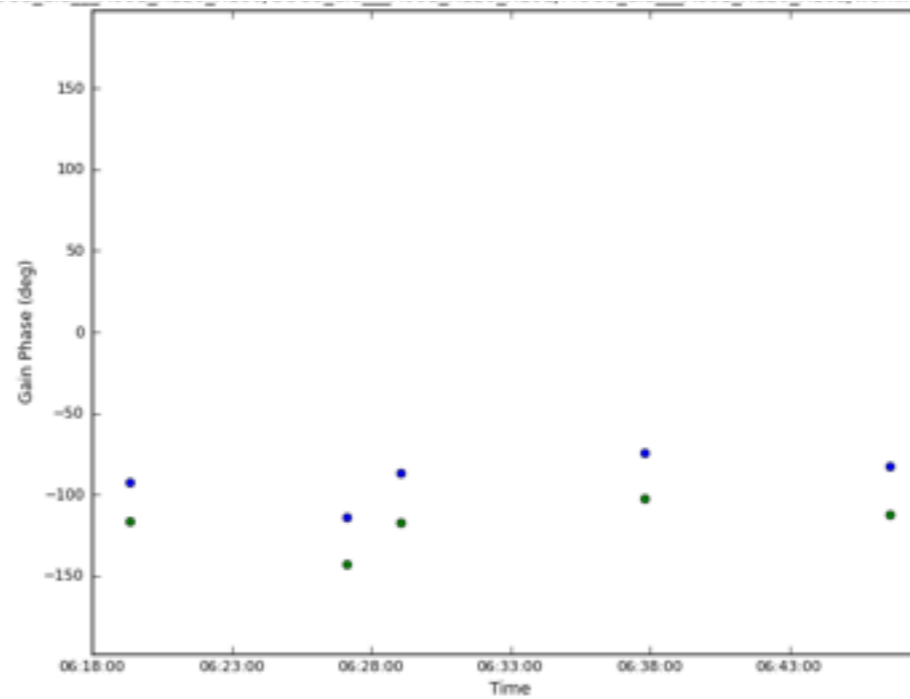
SPW 19,Ant DV19



SPW 21,Ant DV19



SPW 23,Ant DV19



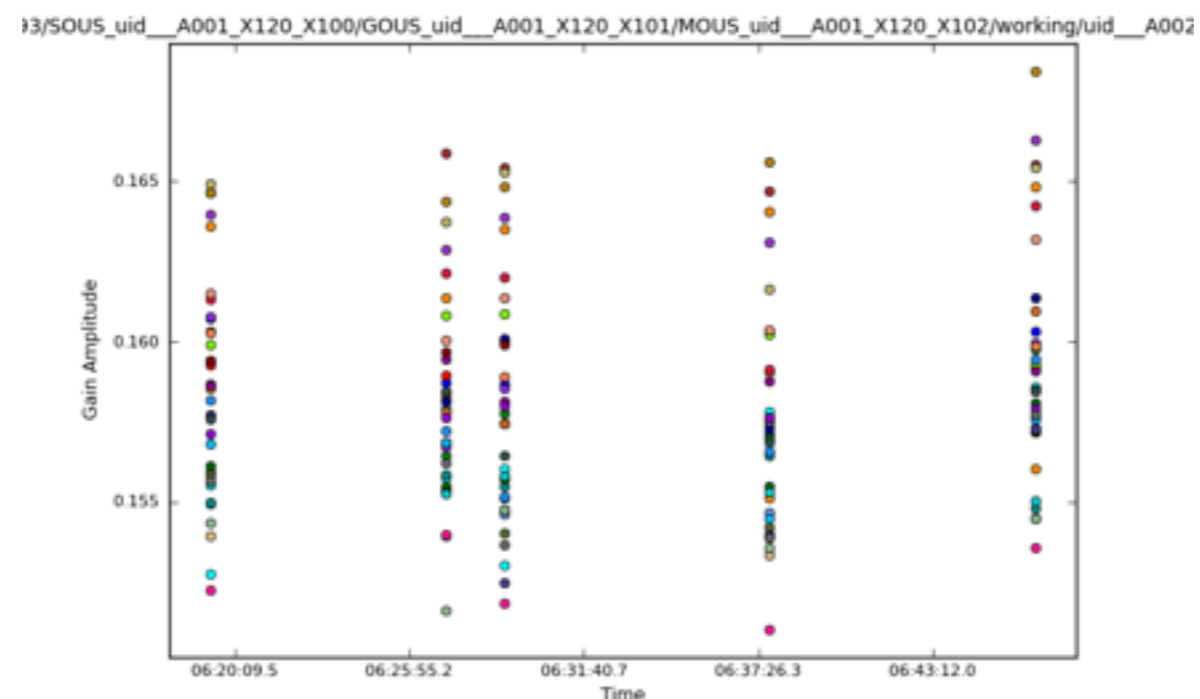
Time ->

# Calibration plots: amplitude vs time

## Step hifa\_timegaincal

The gain amplitude vs. time of the flux, bandpass and phase calibrators, per antenna, per spw or all ant. per spw. Better to see the various spw per antenna plot table `uid*hifa_timegaincal.s | 3_5.spw*solintinf.gacal.tbl` in the `/calibrated/working/` directory

SPW 17, colors=antennas

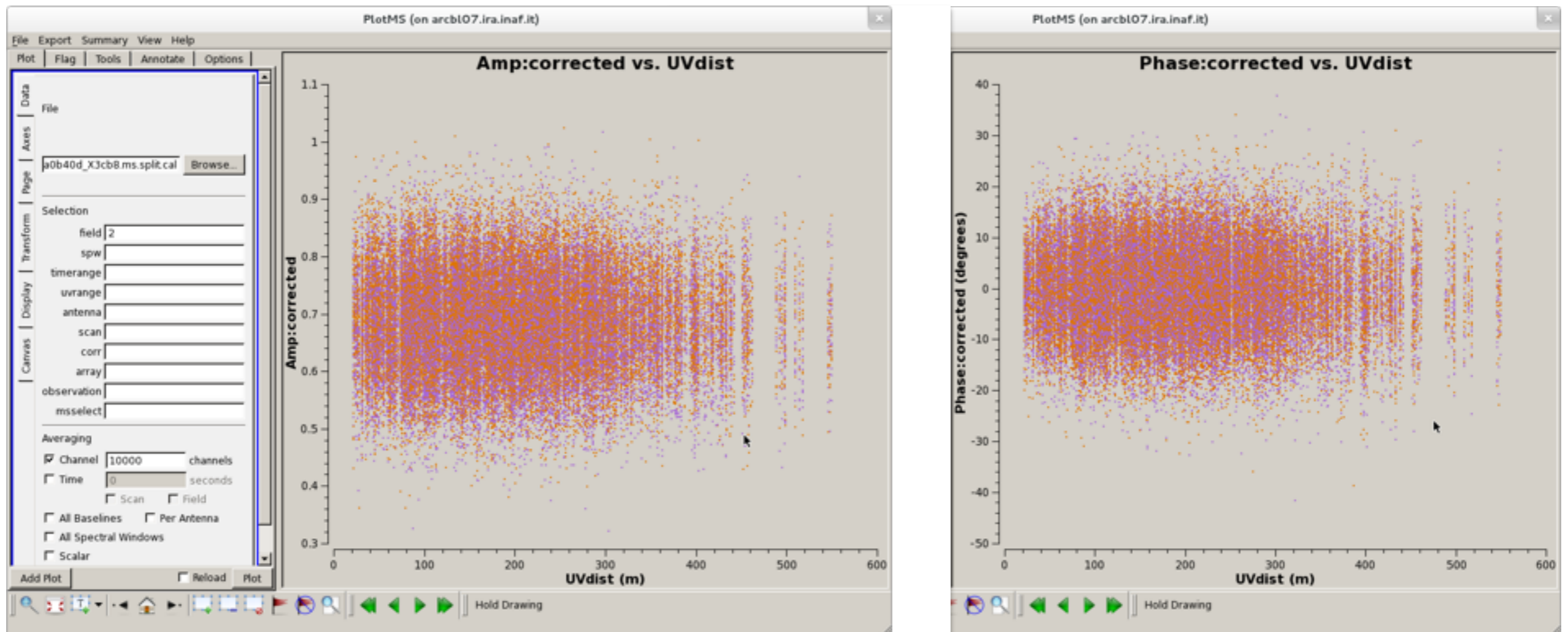


# UVdata phase calibrator

To check whether the phase calibrator is really compact it should have:

- a flat behavior in amplitude vs. uvdistance
- a small scatter ( $\pm 5$ -20 degrees) and flat phases for all uvdistances

These plots are not present in weblog, but can be inspected with plotms()

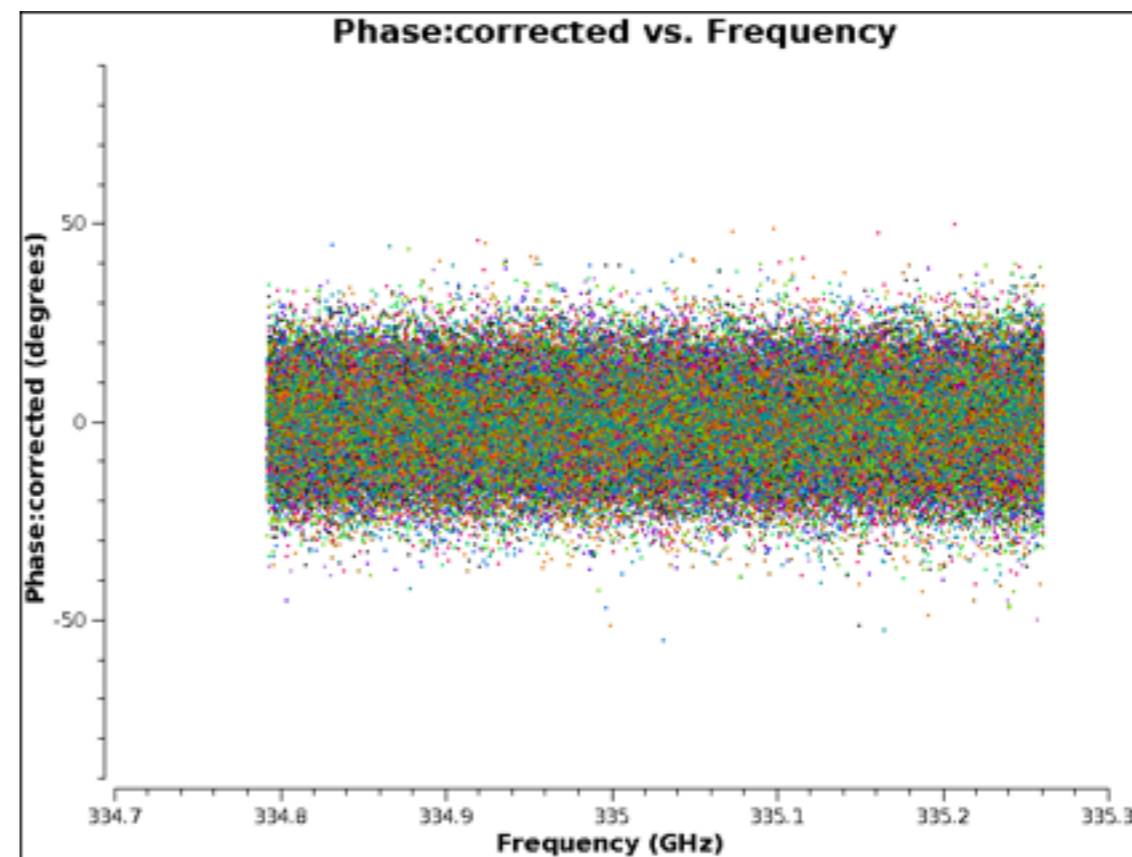
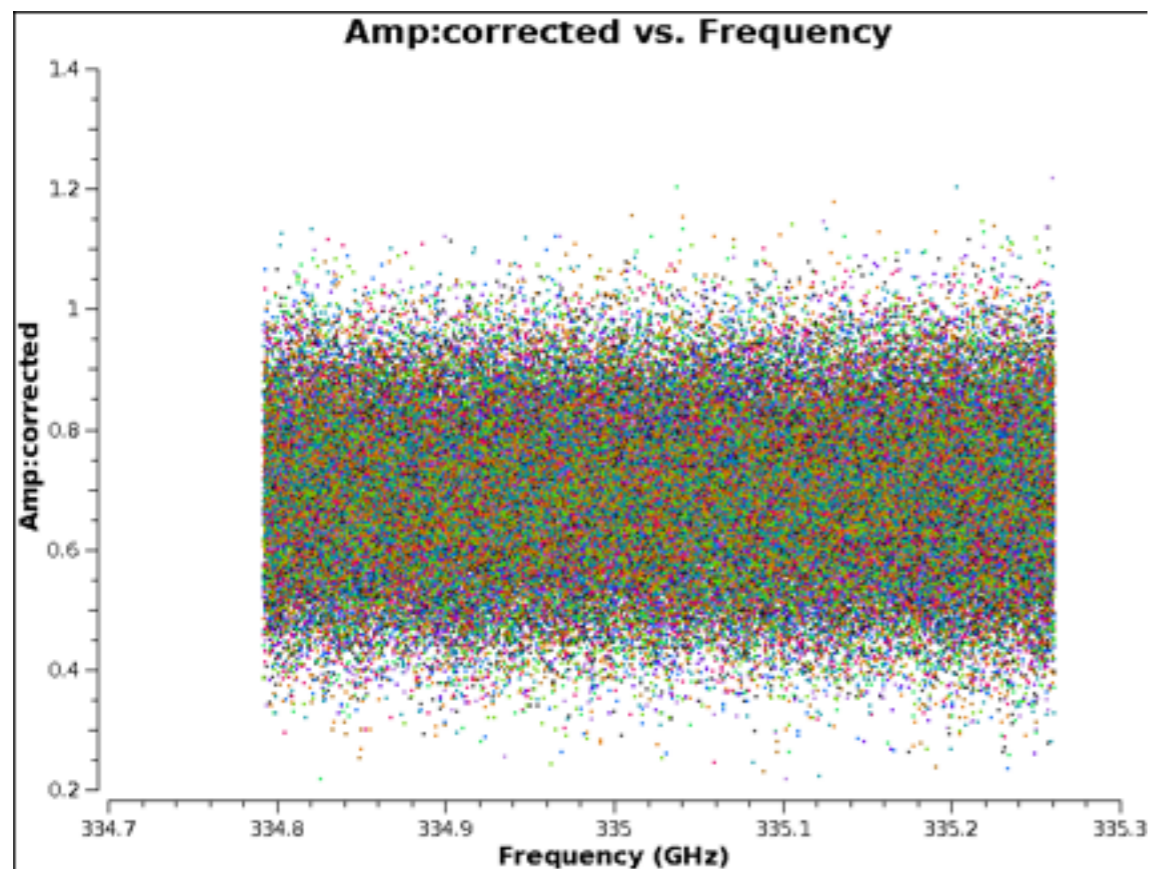




# Freq behavior of phase calibrator

Step `hif_applycal`: plots of phase and bandpass calibrator for each spectral window for amp/phase versus frequency per antenna (colored by antenna) averaged in time.

SPW 19, phase calibrator

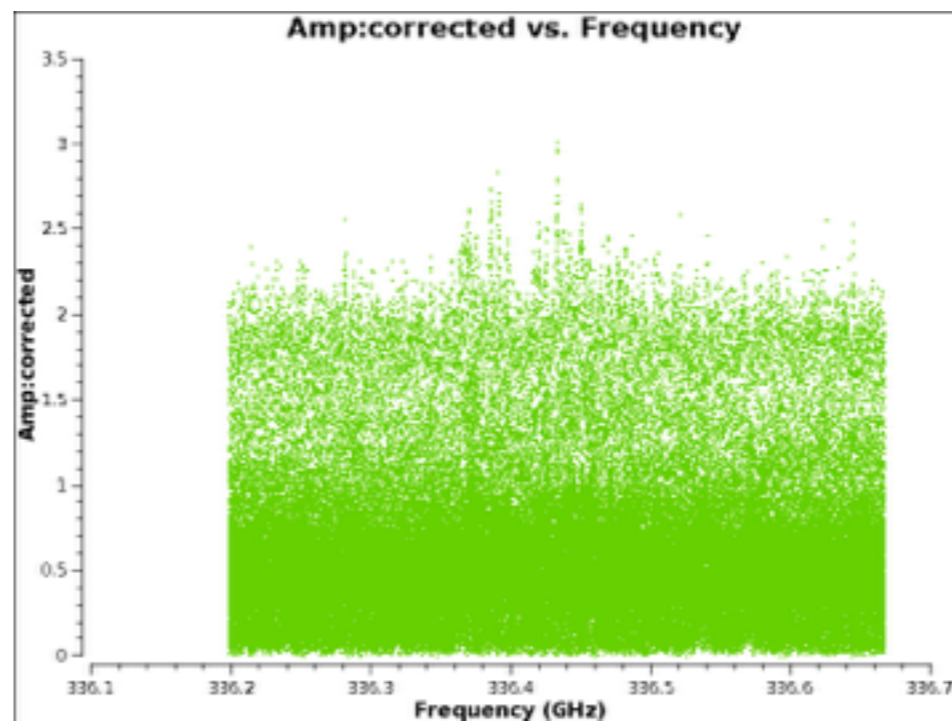
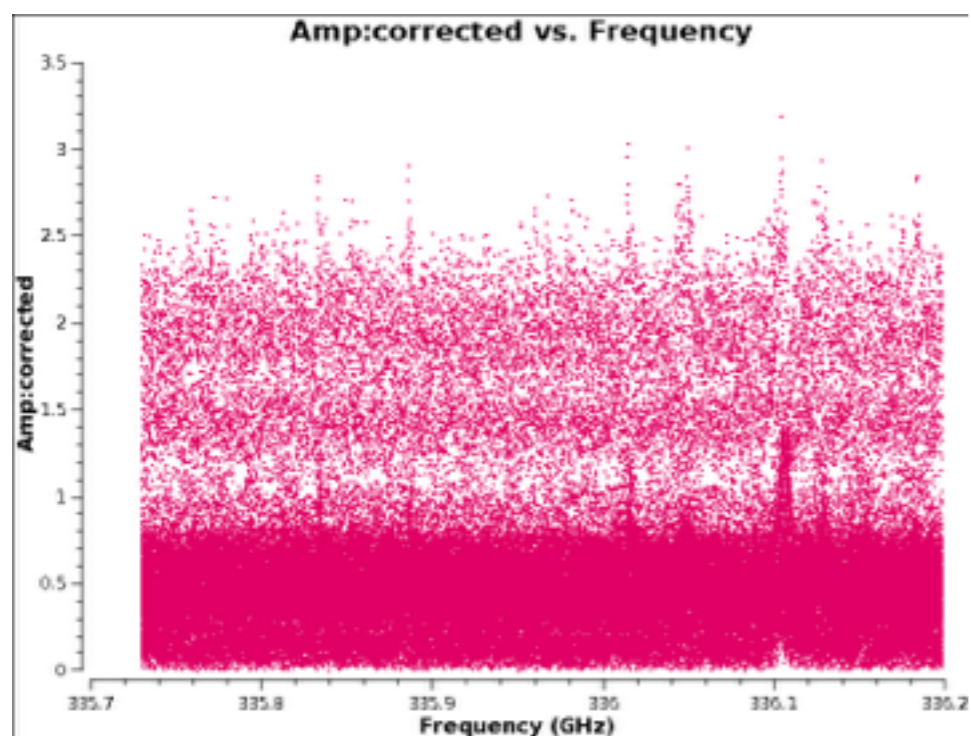
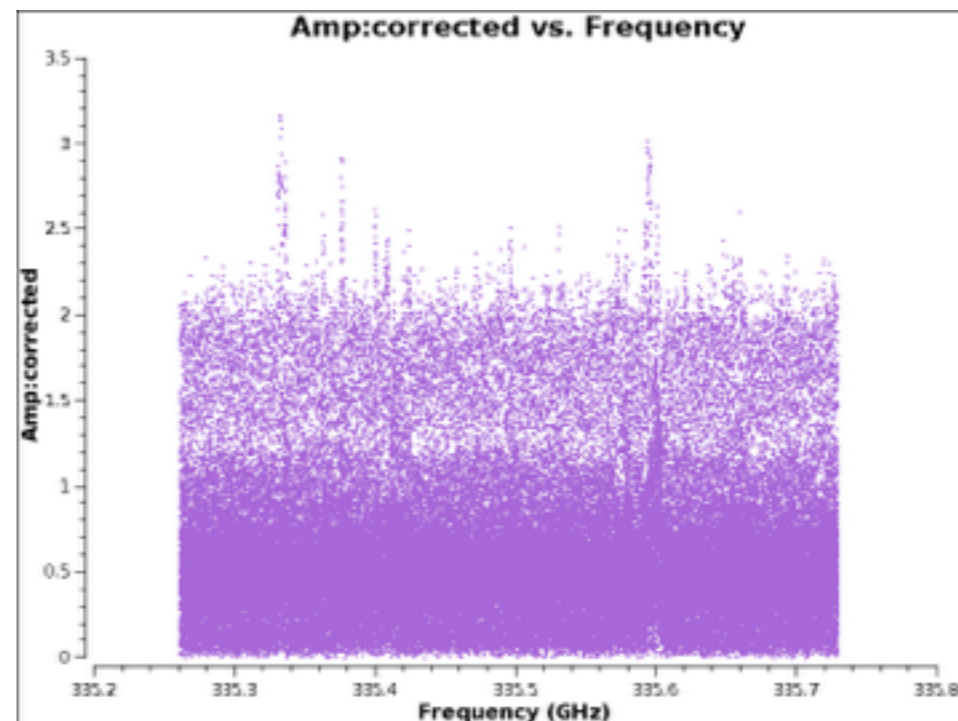
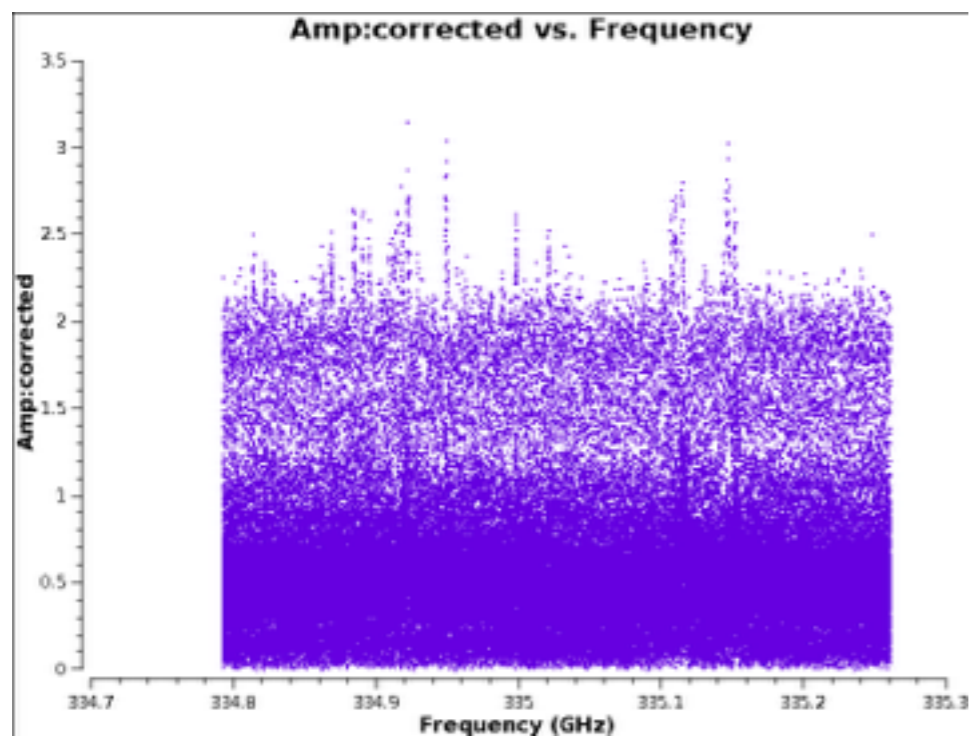






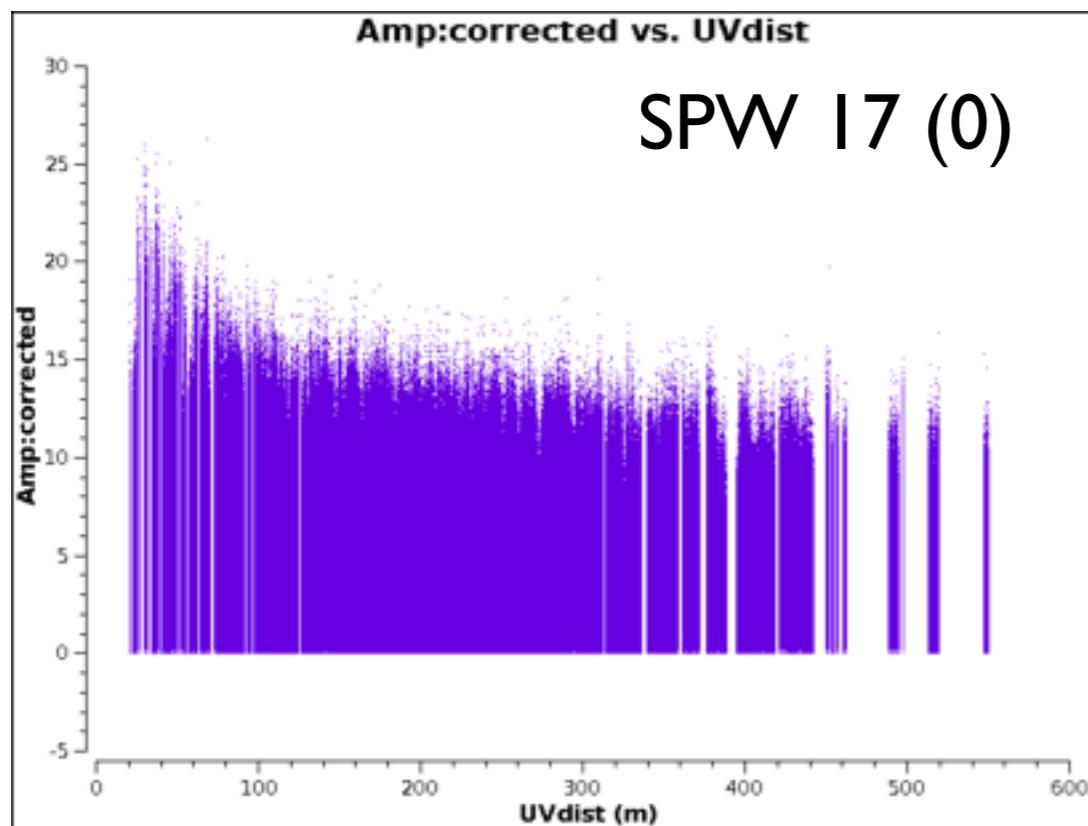
# Target spectra

Step `hif_applycal` gives a first look into target spectra per spw

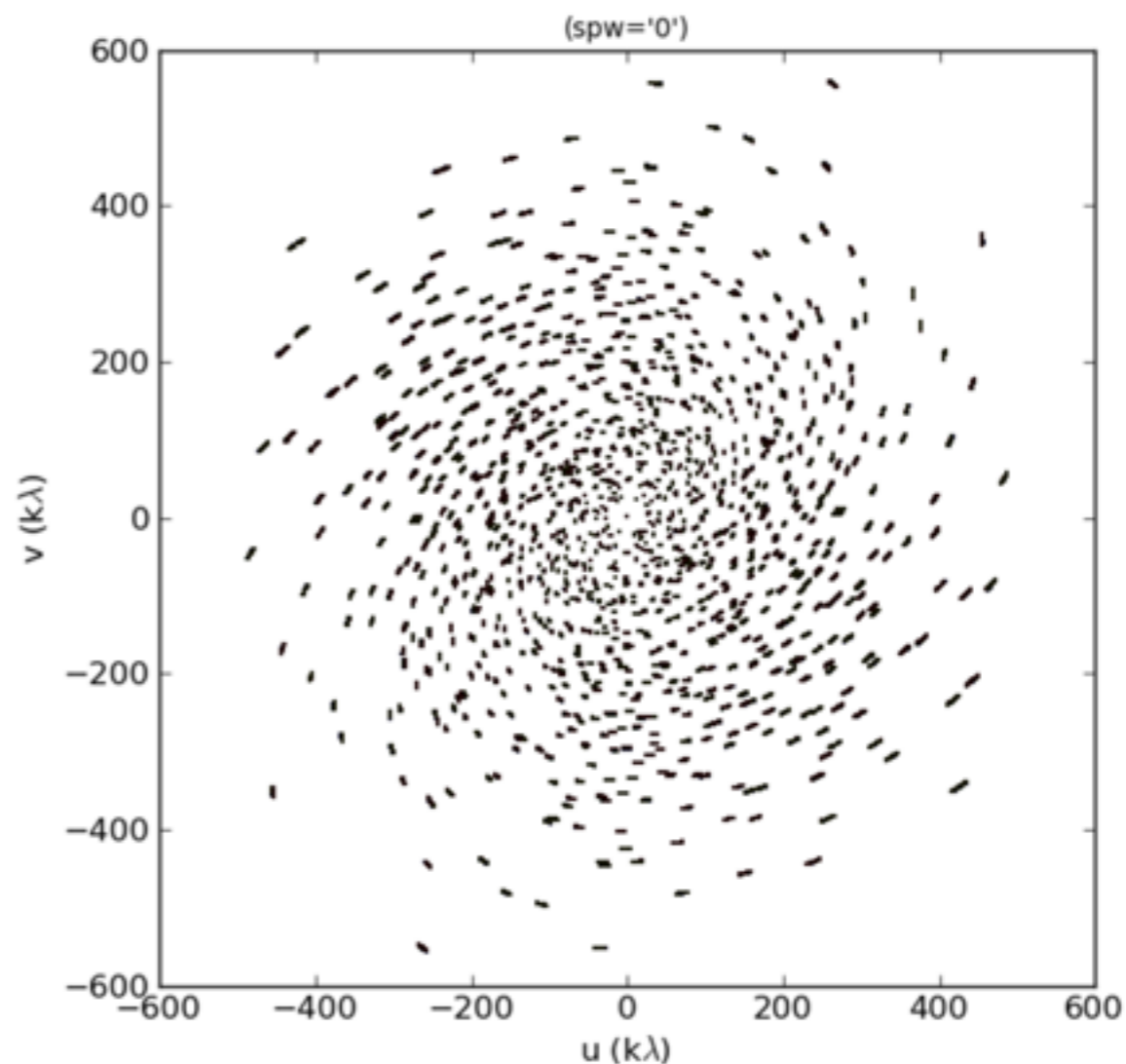


# UVplot of target

Not present in the weblog, but the amplitude vs uvdistance of the target (per spw) can be found in step `hif_applycal`



Uvdist max  $\sim 550\text{m}$  is up to  $\sim 600$   $k\lambda$  (lambda  $0.09\text{cm}$ ), which can be verified with plotms:

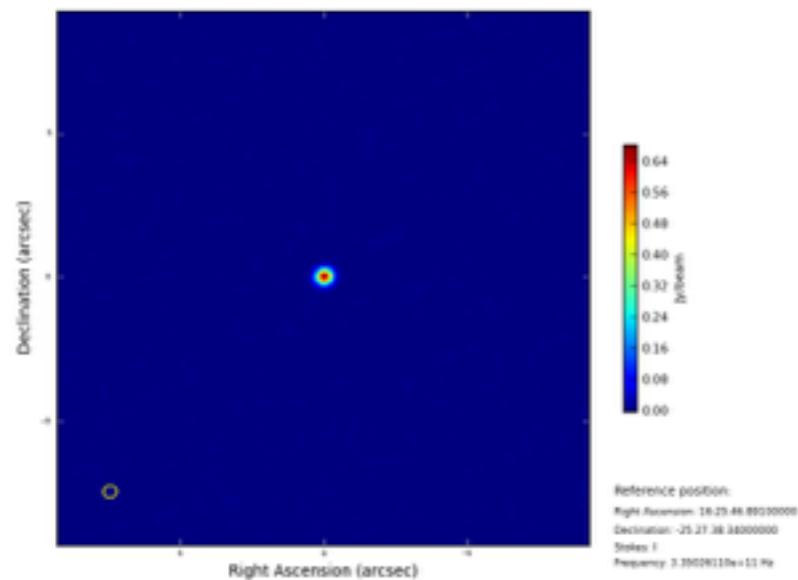


# Phase calibrator image

Step `hif_cleanlist`: images of phase and bandpass calibrators.  
Images of the target have to be made with the `clean` task in Casa.

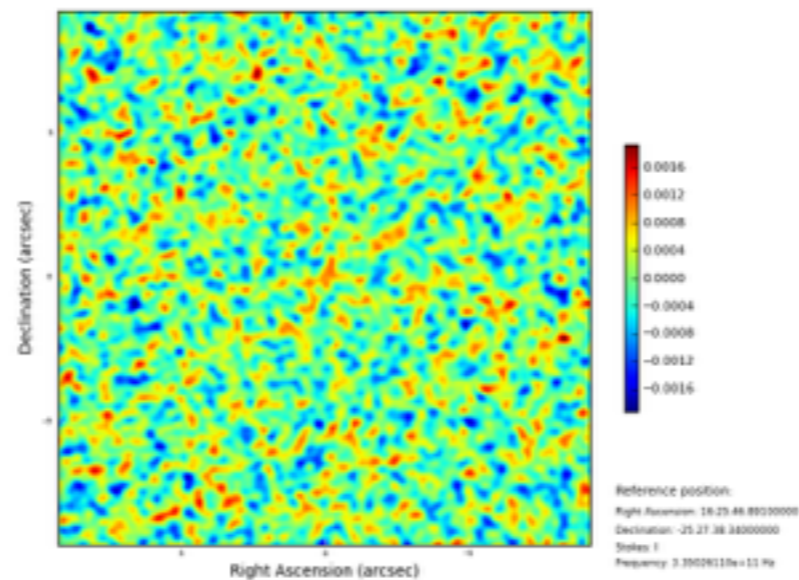
cleaned image

type:image field:j1625-2527 spw:17 iter:1



check on clean:  
residuals

type:residual field:j1625-2527 spw:17 iter:1



dirty beam (PSF)

type:psf field:j1625-2527 spw:17 iter:0

