



Simulating ALMA observations

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Outline

The simulators and cycle 7 proposals

Available simulators

Overview of the CASA simulator

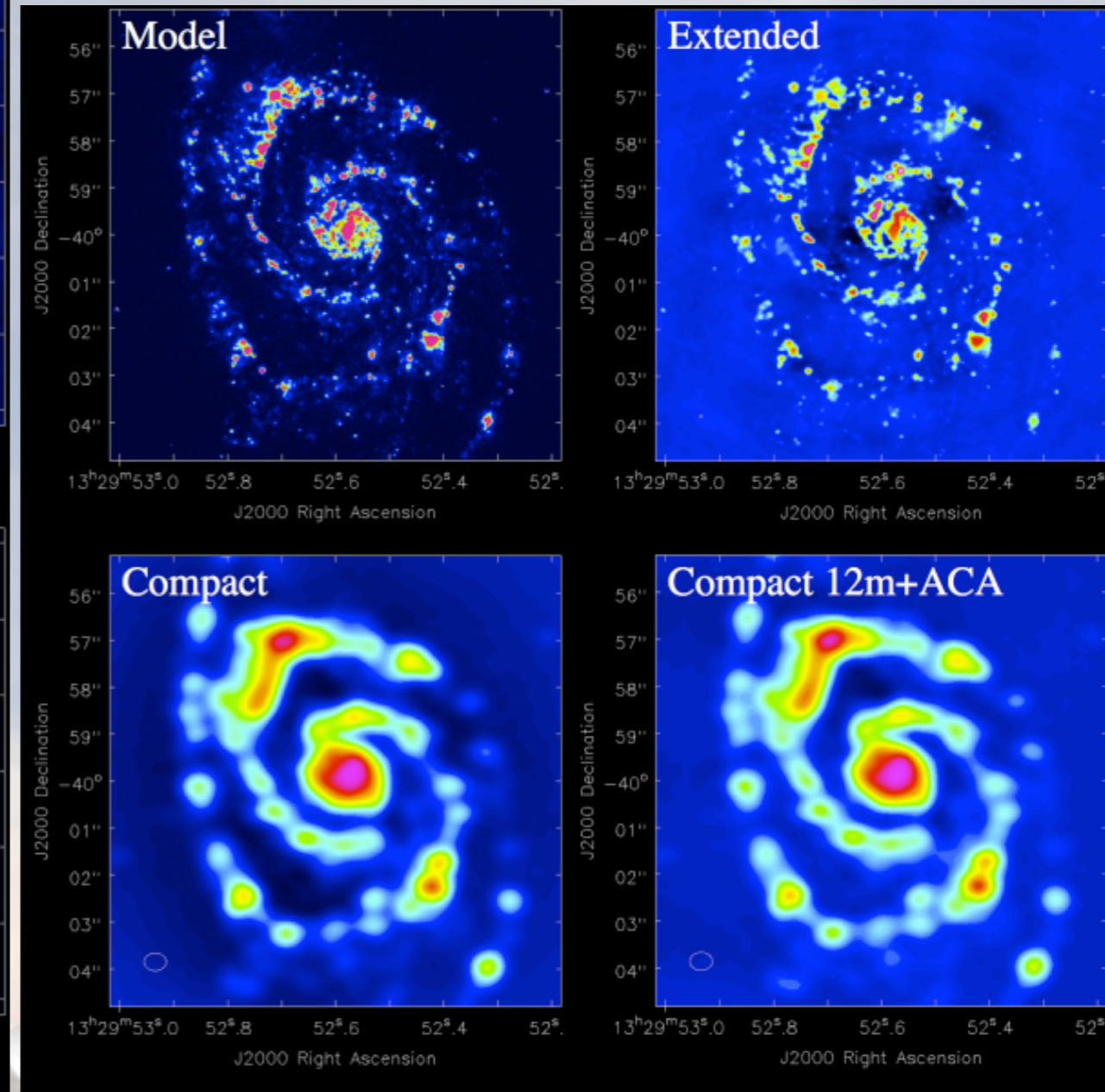
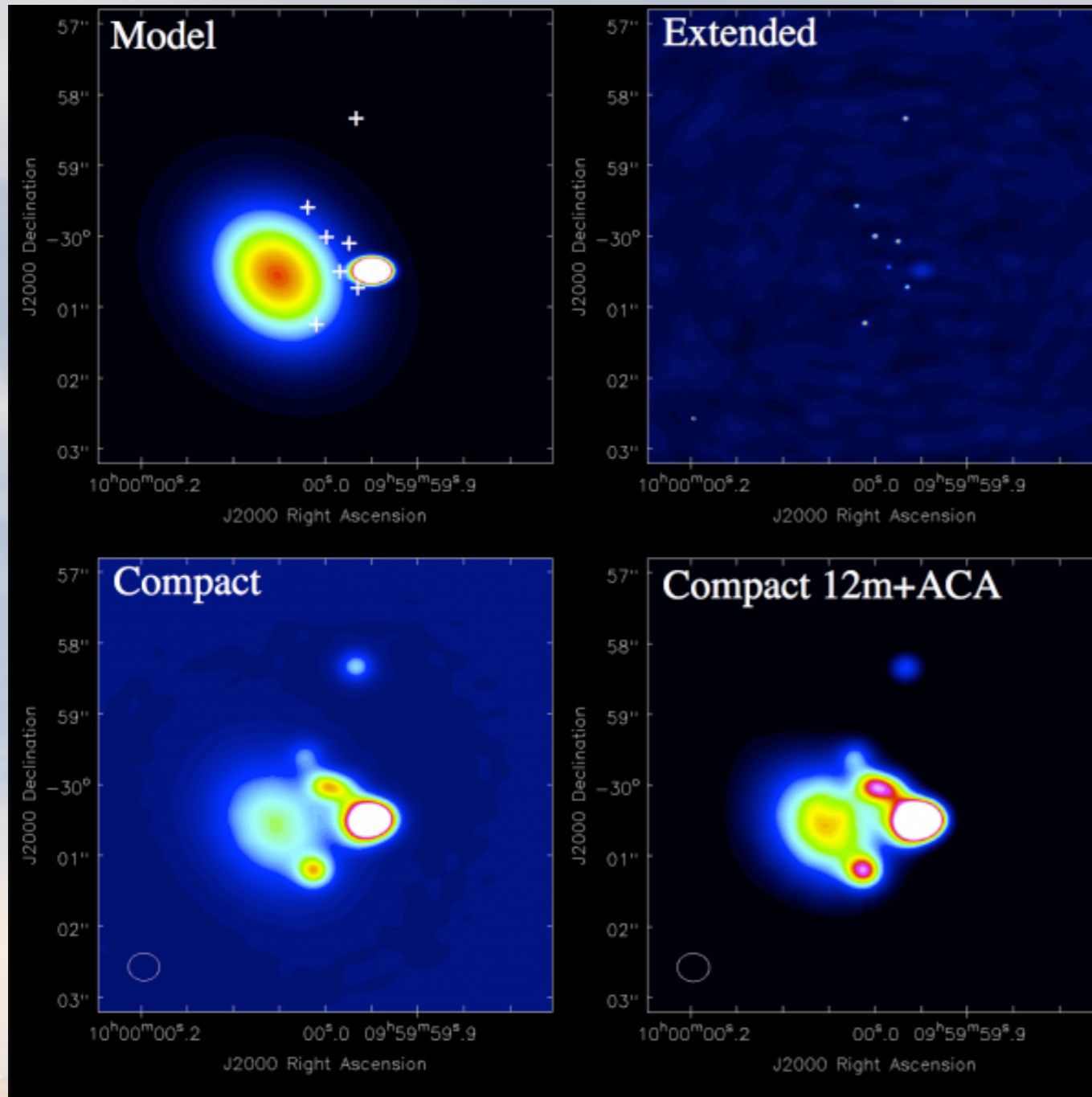


Why should I do an ALMA simulation ?

- You can try out different ALMA configurations and, for a given model, this might help you decide on the required angular resolution and maximum angular scale.
- A simulation can help support the required technical set-up in your proposal
- In particular, it can help you justify the need for complementary ACA/TP or, indeed, ACA-only observations.



Why should I do an ALMA simulation ?



source: CASA simulator pages

Available simulators

CASA simulator

ALMA Observation Support tool

```
# simobserve :: mosaic simulation task:
project          = 'sim'          # root prefix for output file names
skymodel         = '30dor.fits'   # model image to observe
inbright        = '0.06mJy/pixel' # scale surface brightness of brightest pixel e.g. "1.2Jy/pixel"
indirection     = 'J2000 19h00m00 -40d00m00' # set new direction e.g. "J2000 19h00m00 -40d00m00"
incell          = '0.15arcsec'    # set new cell/pixel size e.g. "0.1arcsec"
incenter        = '230GHz'       # set new frequency of center channel e.g. "89GHz" (required even for 2D model)
inwidth         = '2GHz'         # set new channel width e.g. "10MHz" (required even for 2D model)

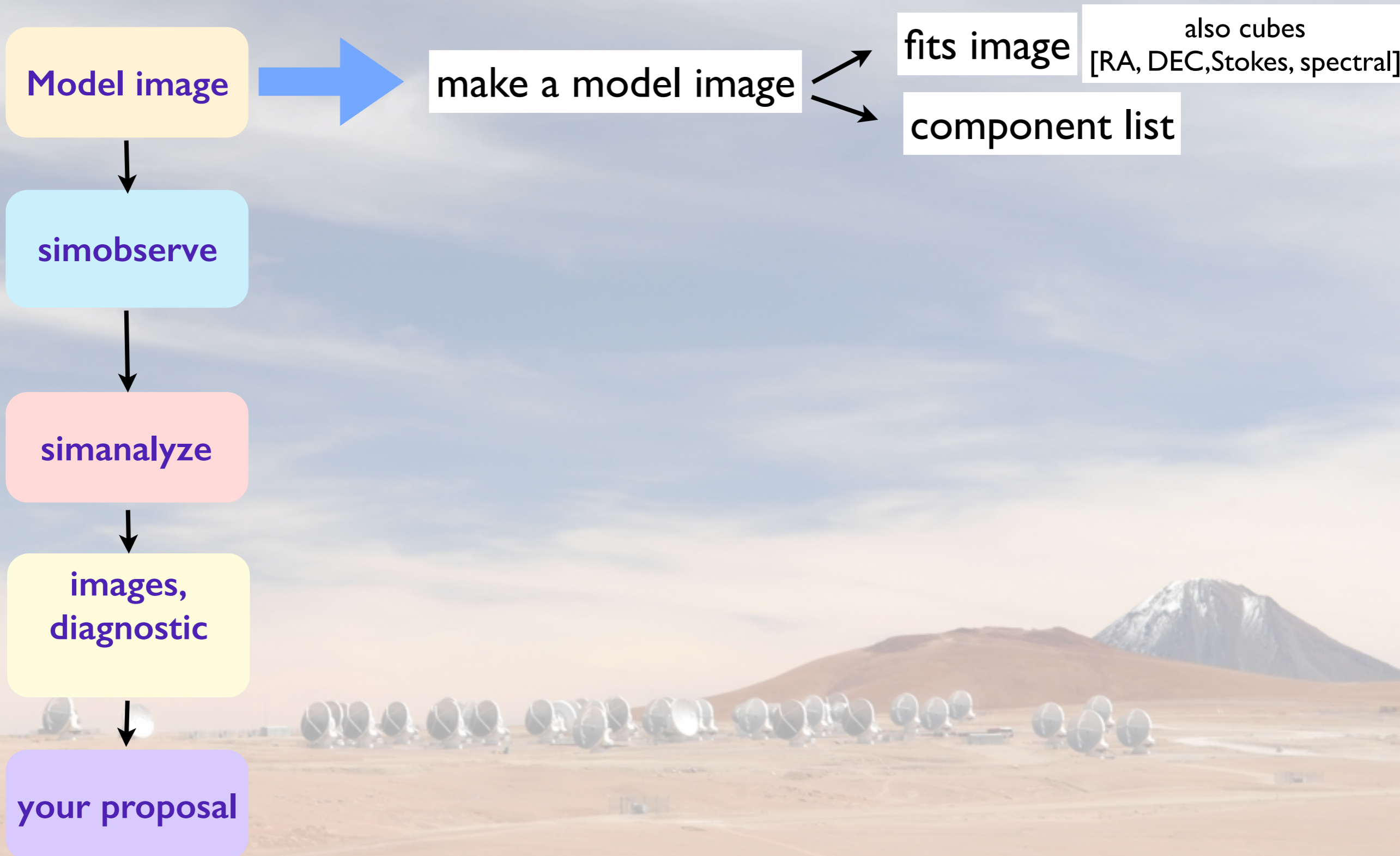
complist        = ''             # componentlist to observe
setpointings   = True           #
  integration    = '600s'        # integration (sampling) time
  direction     = ''             # "J2000 19h00m00 -40d00m00" or "" to center on model
  mapsize       = ['', '']       # angular size of map or "" to cover model
  maptype       = 'topographic'  # hexagonal, square, etc
  pointingspacing = ''           # spacing in between pointings or "0.25PB" or "" for 0.5 PB

obsnode        = 'int'          # observation mode to simulate
  antennalist    = '/usr/lib64/casapy/stable/data/alma/simmos/alma.cycle0.compact.cfg' # interferometer
  antenna position file
  reftime       = '2012/05/21'   # date of observation - not critical unless concatting simulations
  hourangle     = 'transit'      # hour angle of observation center e.g. -3:00:00, or "transit"
  totaltime     = '7200s'       # total time of observation or number of repetitions
  caldirection  = ''            # pt source calibrator [experimental]
  calflux       = '1Jy'

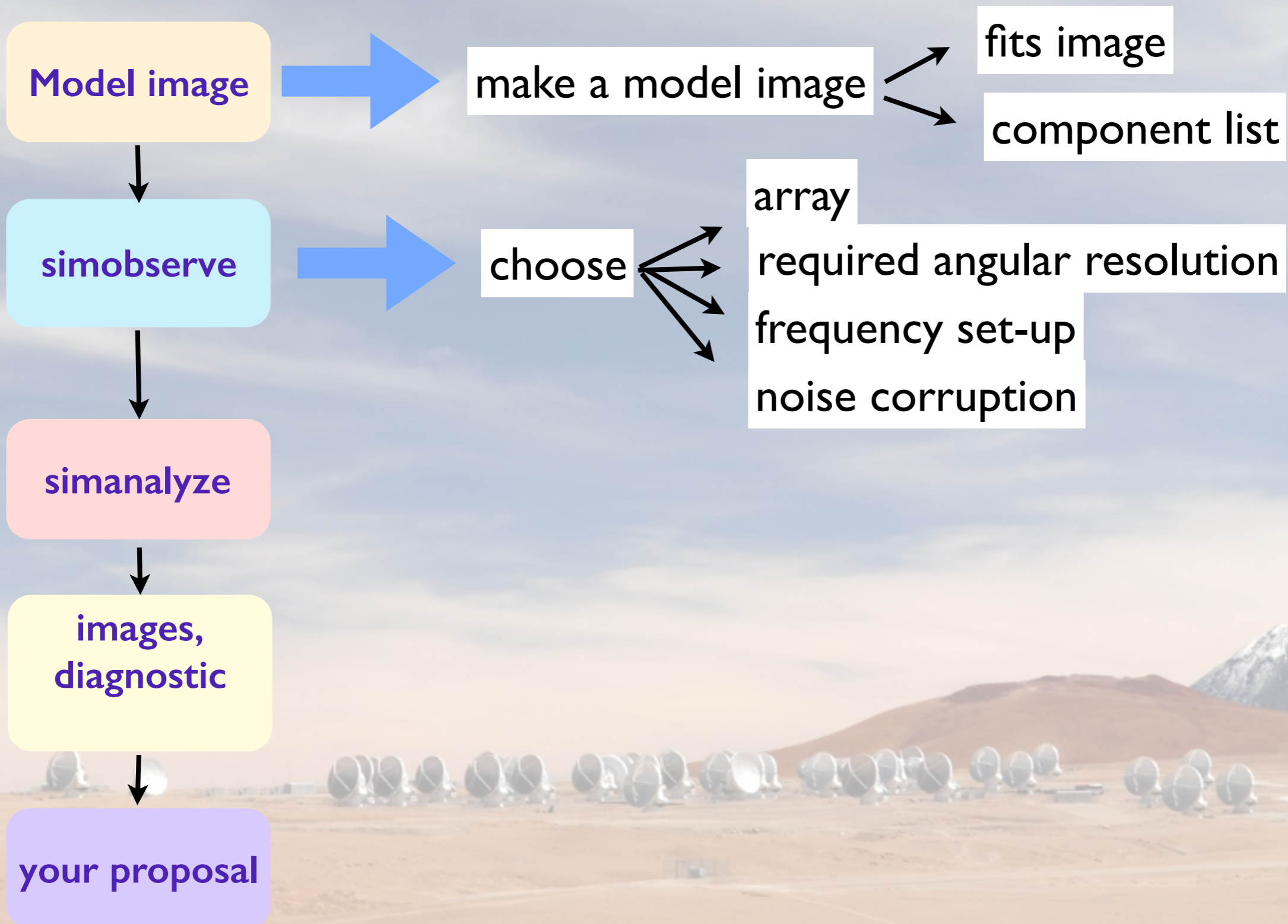
thermalnoise   = ''
leakage        = 0.0
graphics       = 'both'
verbose        = False
overwrite      = True
async          = False
```



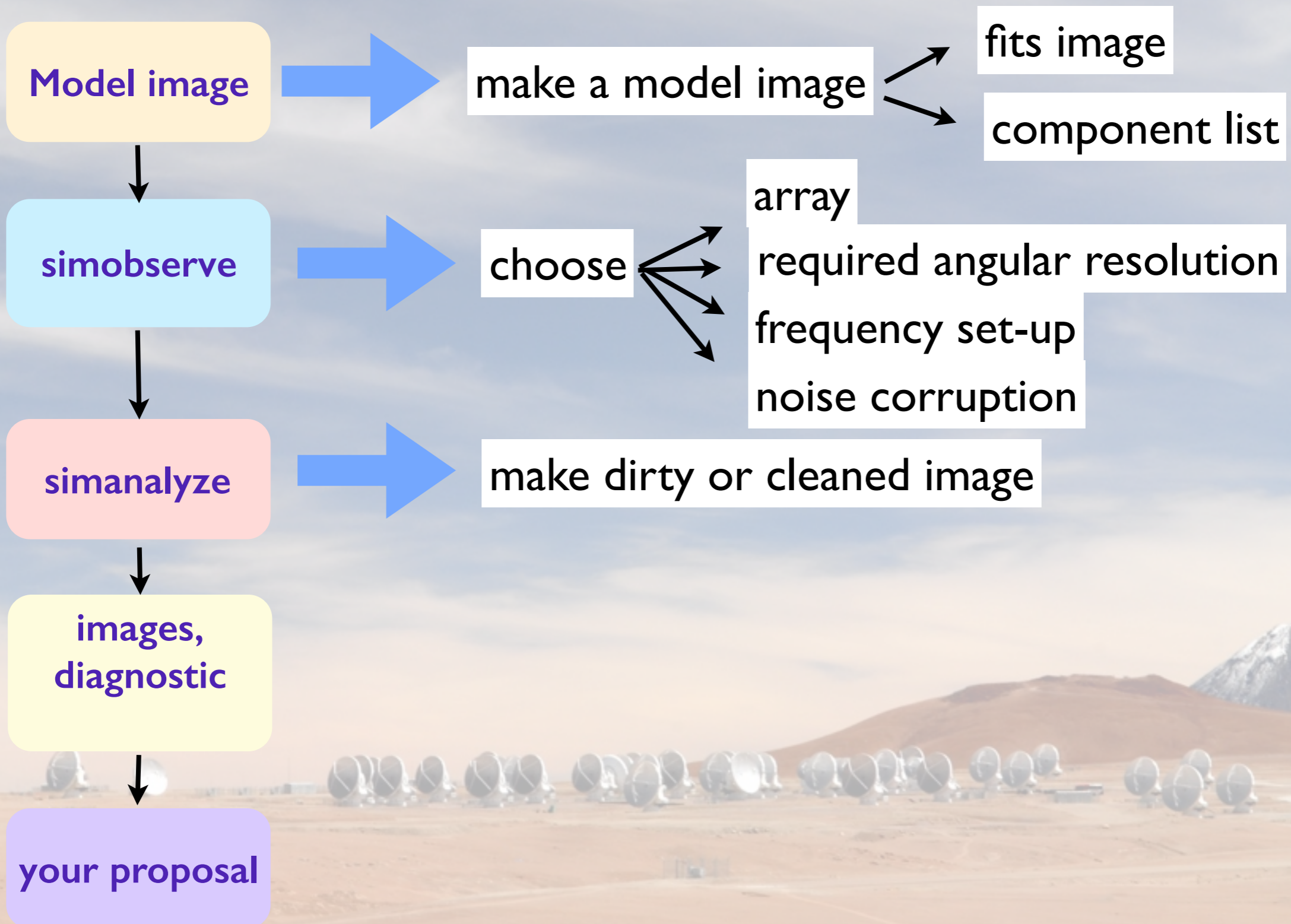
CASA simulator and cycle 7



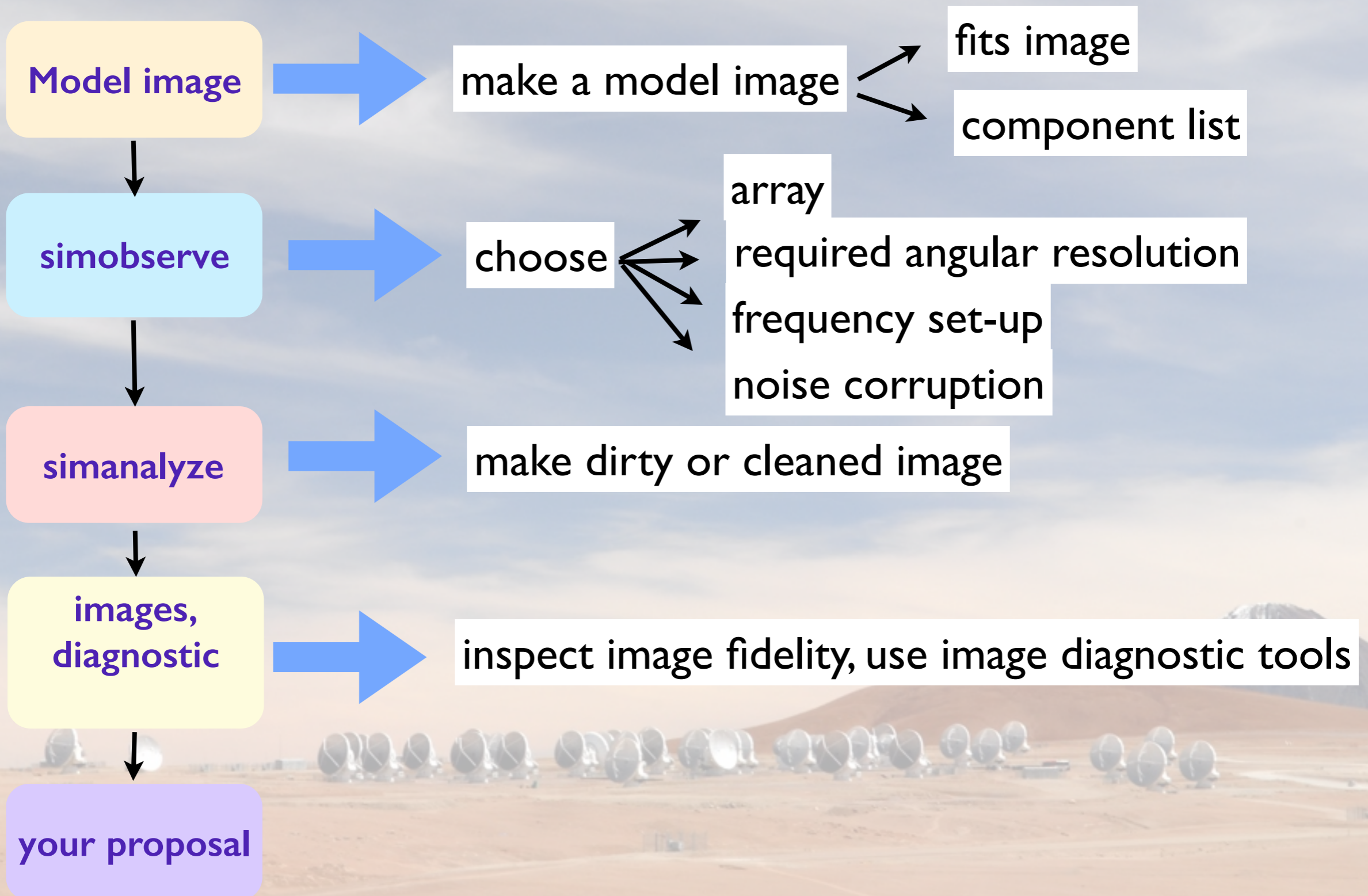
CASA simulator and cycle 7



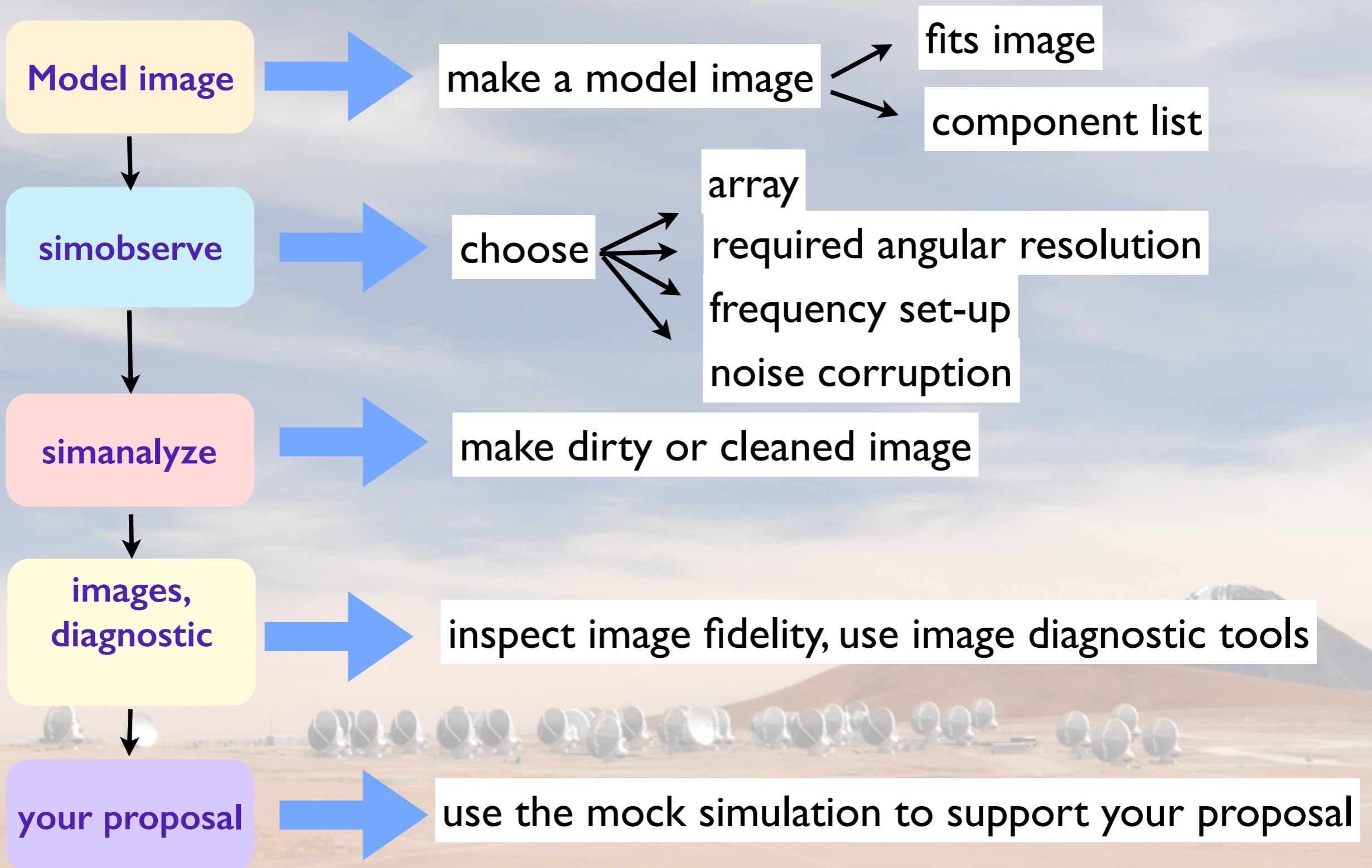
CASA simulator and cycle 7



CASA simulator and cycle 7



CASA simulator and cycle 7



Simobserve

```
# simobserve :: mosaic simulation task:
project          = 'sim'          # root prefix for output file names
skymodel         = '30dor.fits'   # model image to observe
  inbright       = '0.06mJy/pixel' # scale surface brightness of brightest pixel e.g. "1.2Jy/pixel"
  indirection    = 'J2000 10h00m00 -40d00m00' # set new direction e.g. "J2000 19h00m00 -40d00m00"
  incell         = '0.15arcsec'   # set new cell/pixel size e.g. "0.1arcsec"
  incenter       = '230GHz'       # set new frequency of center channel e.g. "89GHz" (required even for
  # model)
  inwidth        = '2GHz'         # set new channel width e.g. "10MHz" (required even for 2D model)

complist         = ''            # componentlist to observe
setpointings    = True          #
  integration     = '600s'       # integration (sampling) time
  direction       = ''           # "J2000 19h00m00 -40d00m00" or "" to center on model
  mapsize         = ['', '']     # angular size of map or "" to cover model
  maptype         = 'topographic' # hexagonal, square, etc
  pointingspacing = ''           # spacing in between pointings or "0.25PB" or "" for 0.5 PB

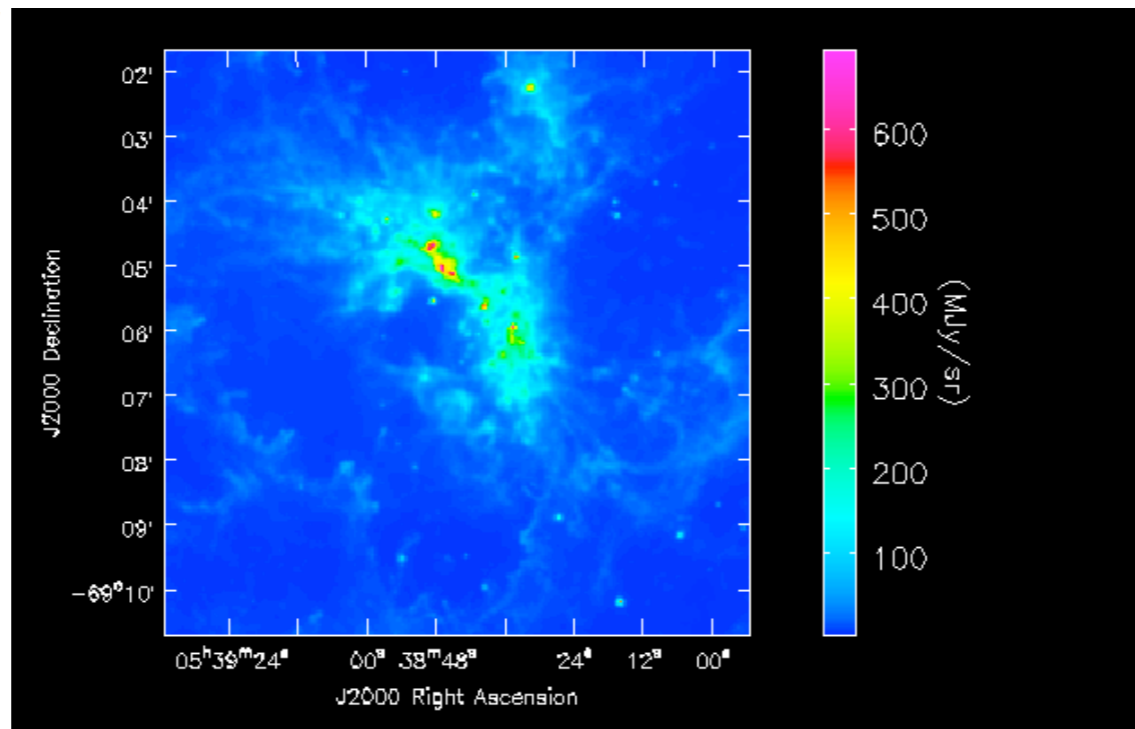
obsmode         = 'int'         # observation mode to simulate
  # [int(interferometer)|sd(singledish)|""(none)]
  antennalist     = '/usr/lib64/casapy/stable/data/alma/simmos/alma.cycle0.compact.cfg' # interferometer
  # antenna position file
  refdate         = '2012/05/21' # date of observation - not critical unless concatting simulations
  hourangle       = 'transit'    # hour angle of observation center e.g. -3:00:00, or "transit"
  totaltime       = '7200s'     # total time of observation or number of repetitions
  caldirection    = ''           # pt source calibrator [experimental]
  calflux         = '1Jy'

thermalnoise    = ''            # add thermal noise: [tsys-atm|tsys-manual|""]
leakage         = 0.0            # cross polarization (interferometer only)
graphics        = 'both'        # display graphics at each stage to [screen|file|both|none]
verbose         = False         #
overwrite       = True          # overwrite files starting with $project
async           = False         # If true the taskname must be started using simobserve(...)
```

Simobserve

```
# simobserve :: mosaic simulation task:
project          = 'sim'          # root prefix for
skymodel         = '30dor.fits'   # model image to
  inbright       = '0.06mJy/pixel' # scale surface
  indirection    = 'J2000 10h00m00 -40d00m00' # set
  incell         = '0.15arcsec'   # set new cell
  incenter       = '230GHz'       # set new freq
  inwidth        = '2GHz'         # set new chan
complist         = ''             # component list
```

Model
image



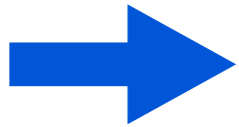
Spitzer IRAC 89 micron
image of 30 Doradus

Simobserve

interferometer (int) or single-dish (sd)

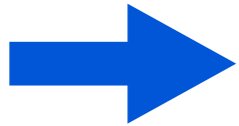


```
obsmode = 'int' # observation mode to simulate
           # [int(interferometer)|sd(singledish)|""(none)]
antennalist = '/usr/lib64/casapy/stable/data/alma/simmos/alma.cycle0.compact.cfg'
```



Cycle 7 representative configurations files

<https://almascience.eso.org/documents-and-tools/cycle7/alma-configuration-files>



or, give the required angular resolution

"alma;0.5arcsec"

(only works for full ALMA !)

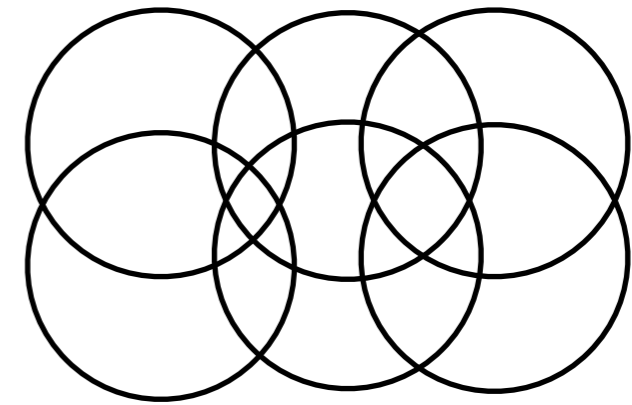


Simobserve

```
incenter = '230GHz'  
inwidth = '2GHz'
```

spectral set-up

```
complist = ''  
setpointings = True  
integration = '600s'  
direction = ''  
mapsize = ['', '']  
maptype = rectangle or hexagonal  
pointingspacing = ''
```



→ the simulator allows you to
test
several pointing scenarios

observing
time and
pointing

Simobserve

```
# simobserve :: mosaic simulation task:
```

```
project = 'sim'
```

```
skymodel = '30dor.fits'
```

```
inbright = '0.06mJy/pixel'
```

```
indirection = 'J2000 10h00m00 -40d00m00'
```

```
incell = '0.15arcsec'
```

```
incenter = '230GHz'
```

```
inwidth = '2GHz'
```

```
complist = ''
```

```
setpointings = True
```

```
integration = '600s'
```

```
direction = ''
```

```
mapsize = ['', '']
```

```
maptype = 'topographic'
```

```
pointingspacing = ''
```

```
obsmode = 'int'
```

```
antennalist = '/usr/lib64/casapy/stable/data/alma/simmos/alma.cycle0.compact.cfg' # inter
```

```
refdate = '2012/05/21'
```

```
hourangle = 'transit'
```

```
totaltime = '7200s'
```

```
caldirection = ''
```

```
calflux = '1Jy'
```

```
thermalnoise = 'tsys-atm'
```

```
user_pwv = 1.0
```

```
t_ground = 269.0
```

```
seed = 11111
```

```
leakage = 0.0
```

```
graphics = 'both'
```

```
verbose = False
```

```
overwrite = True
```

```
async = False
```

model image

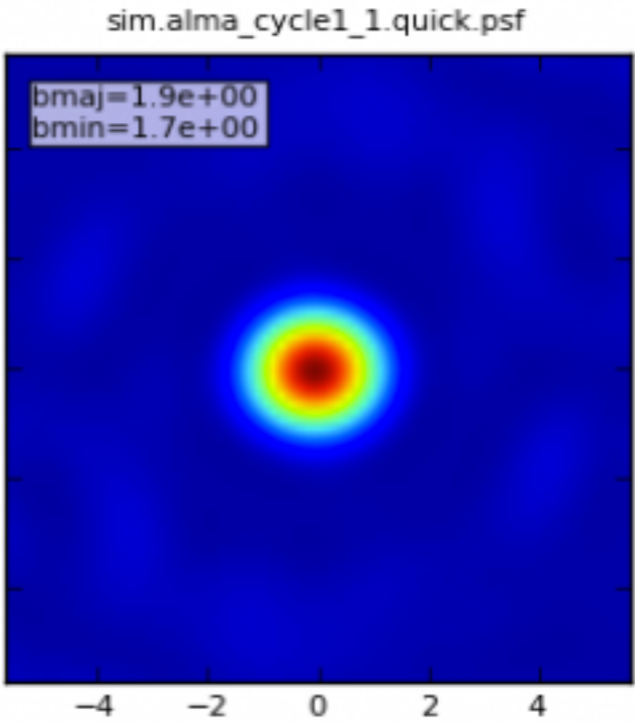
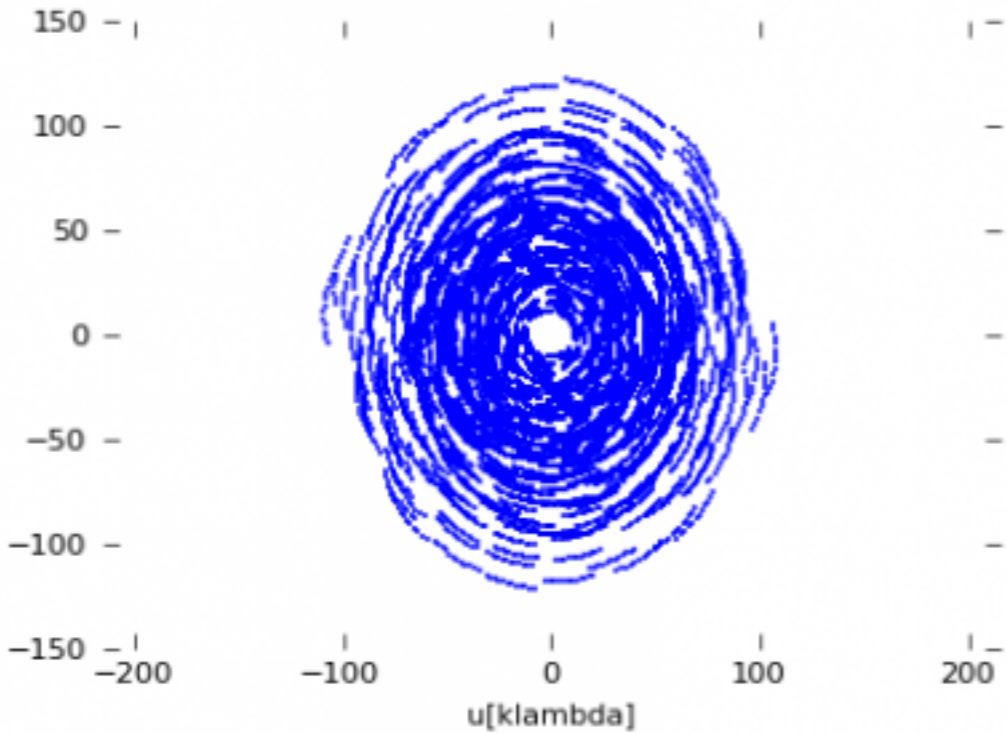
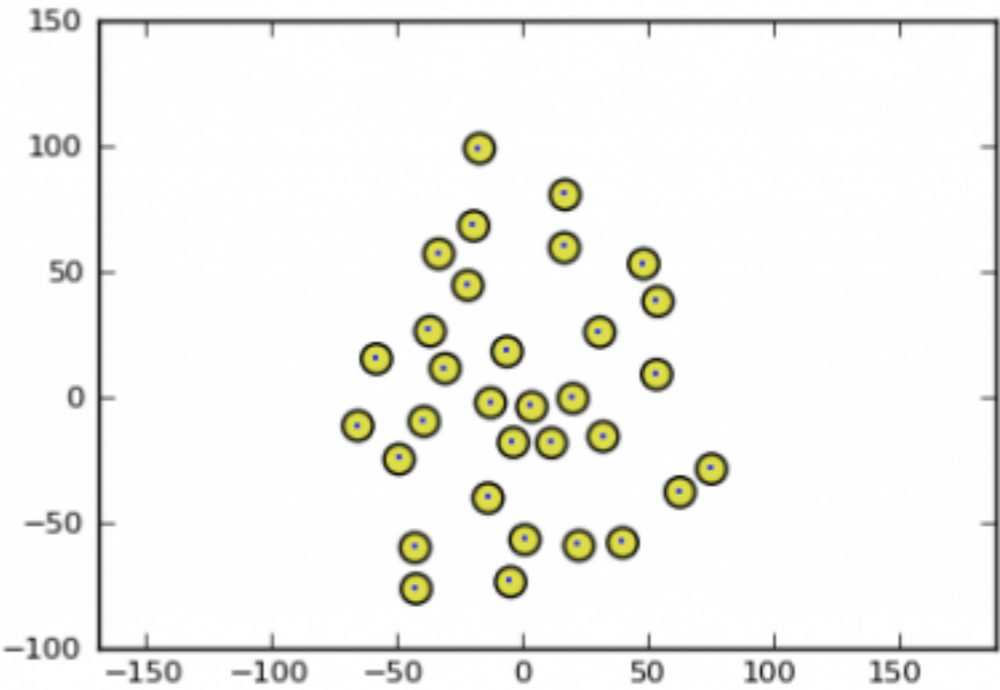
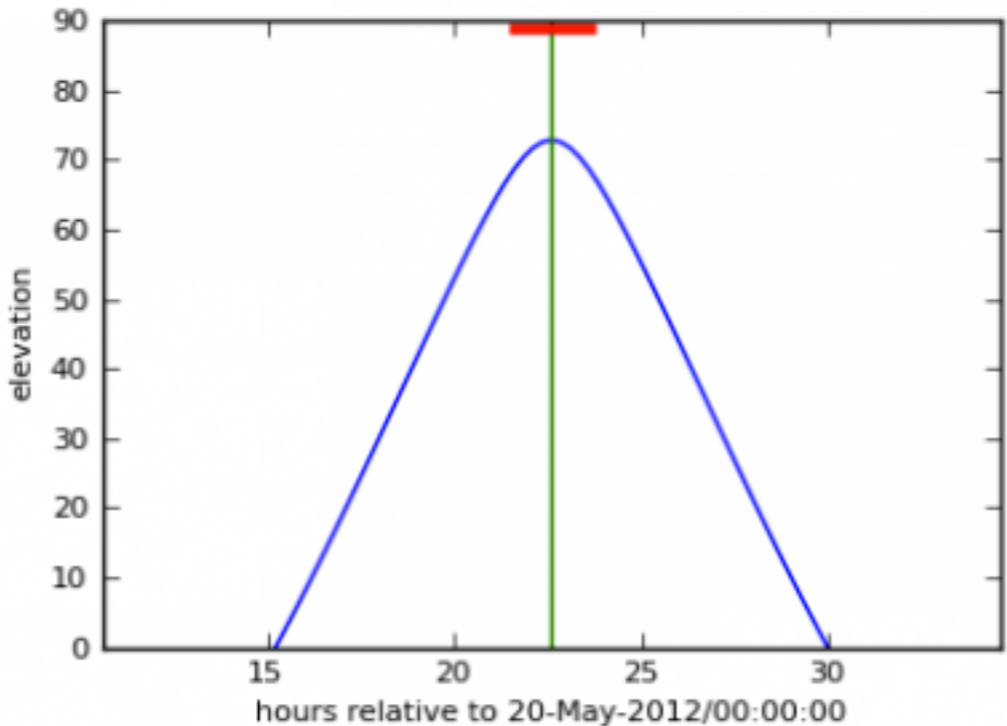
spectral set-up

observing
time and
pointing

array choice
or
required
angular
resolution

noise
addition

Simobserve



Imaging step

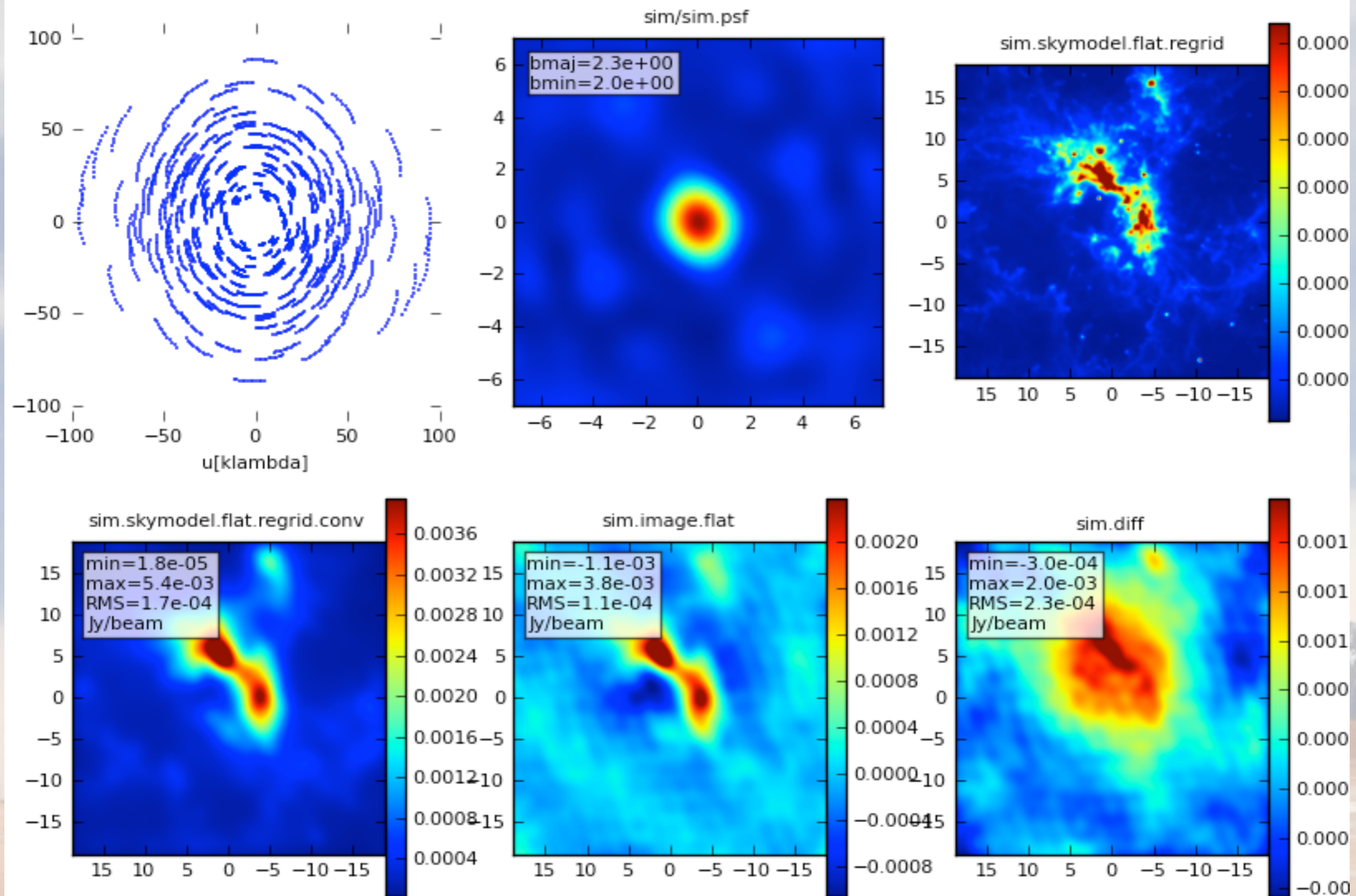
simanalyze

```
# simanalyze :: image and analyze simulated datasets
project          = 'sim'          # root prefix for output file names
image            = True           # (re)image $project.*.ms to $project.image
  vis            = 'default'     # Measurement Set(s) to image
  modelimage     = ''            # prior image to use in clean e.g. existing single dish image
  imsize         = 0             # output image size in pixels (x,y) or 0 to match model
  imdirection   = ''            # set output image direction, (otherwise center on the model)
  cell           = ''            # cell size with units or "" to equal model
  niter          = 500           # maximum number of iterations (0 for dirty image)
  threshold      = '0.1mJy'     # flux level (+units) to stop cleaning
  weighting      = 'natural'     # weighting to apply to visibilities
  mask           = []           # Cleanbox(es), mask image(s), region(s), or a level
  outertaper     = []           # uv-taper on outer baselines in uv-plane
  stokes         = 'I'          # Stokes params to image

analyze          = True          # (only first 6 selected outputs will be displayed)
  showuv         = True          # display uv coverage
  showpsf        = True          # display synthesized (dirty) beam (ignored in single dish simulation)
  showmodel      = True          # display sky model at original resolution
  showconvolved  = False         # display sky model convolved with output beam
  showclean      = True          # display the synthesized image
  showresidual   = False         # display the clean residual image (ignored in single dish simulation)
  showdifference = True          # display difference image
  showfidelity   = True          # display fidelity

graphics         = 'both'       # display graphics at each stage to [screen|file|both|none]
verbose          = False
overwrite        = True         # overwrite files starting with $project
async           = False         # If true the taskname must be started using simanalyze(...)
```

Simanalyze imaging output

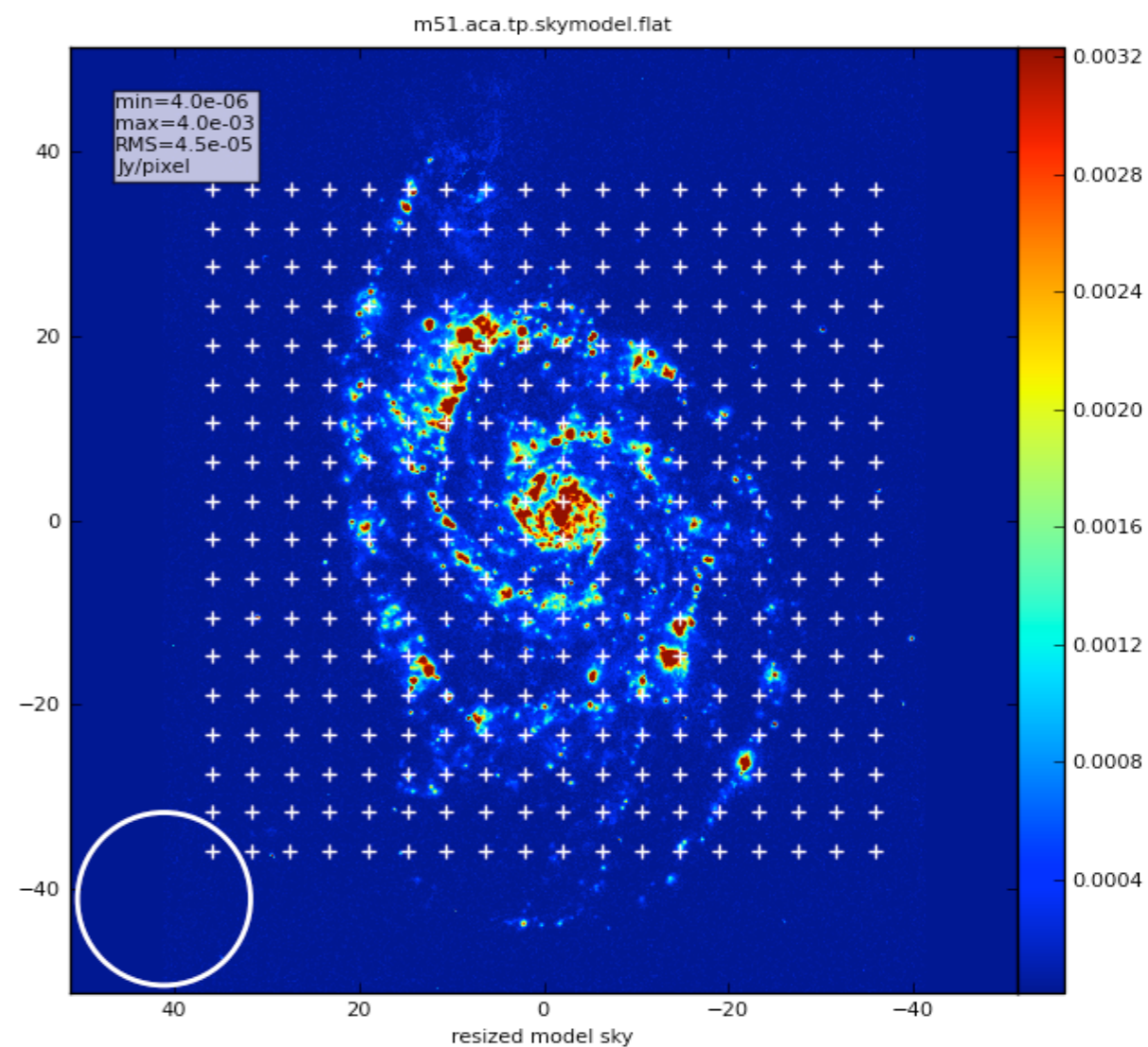
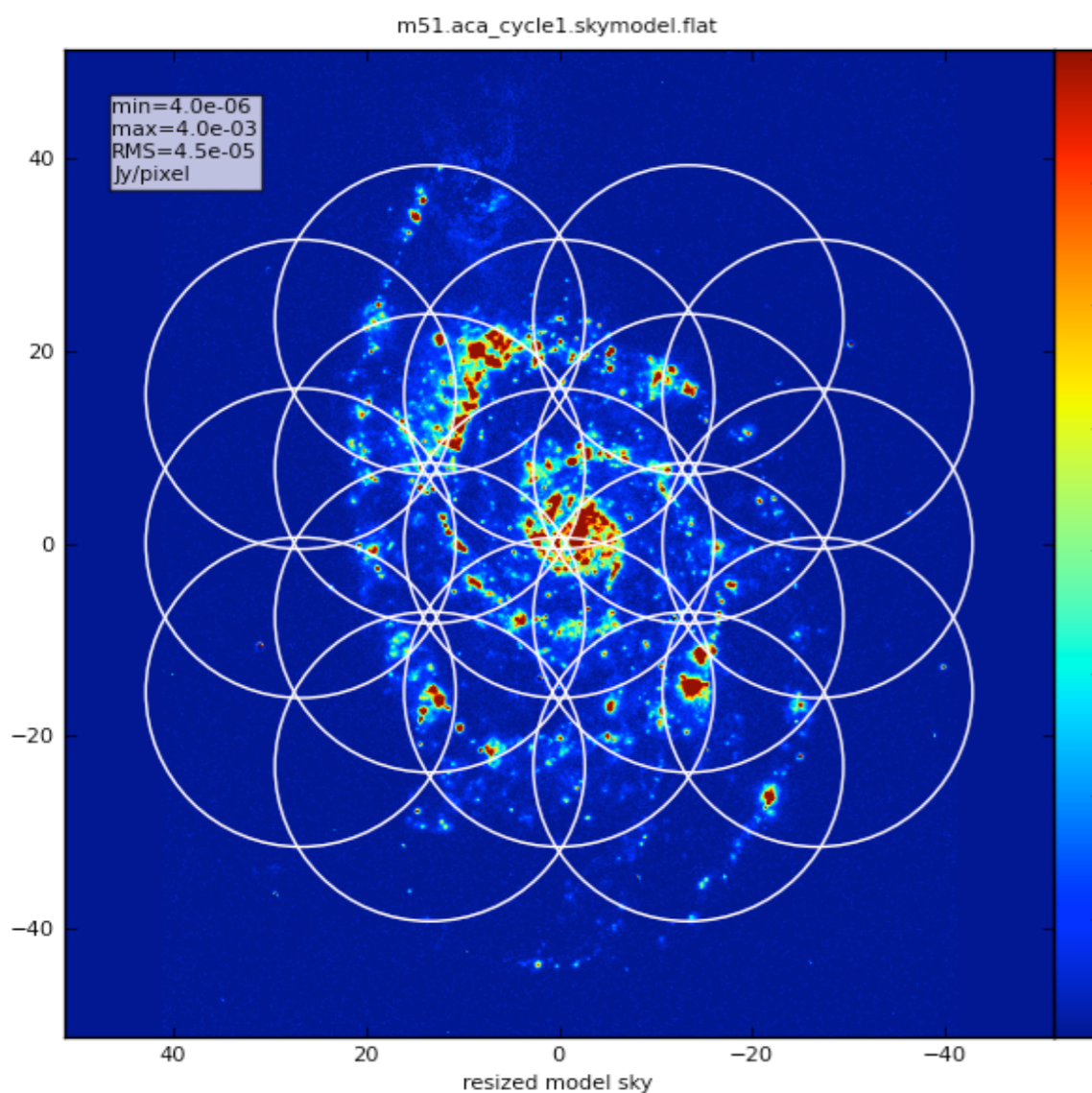


Simalma - combination of SD/ACA/ALMA during de-convolution

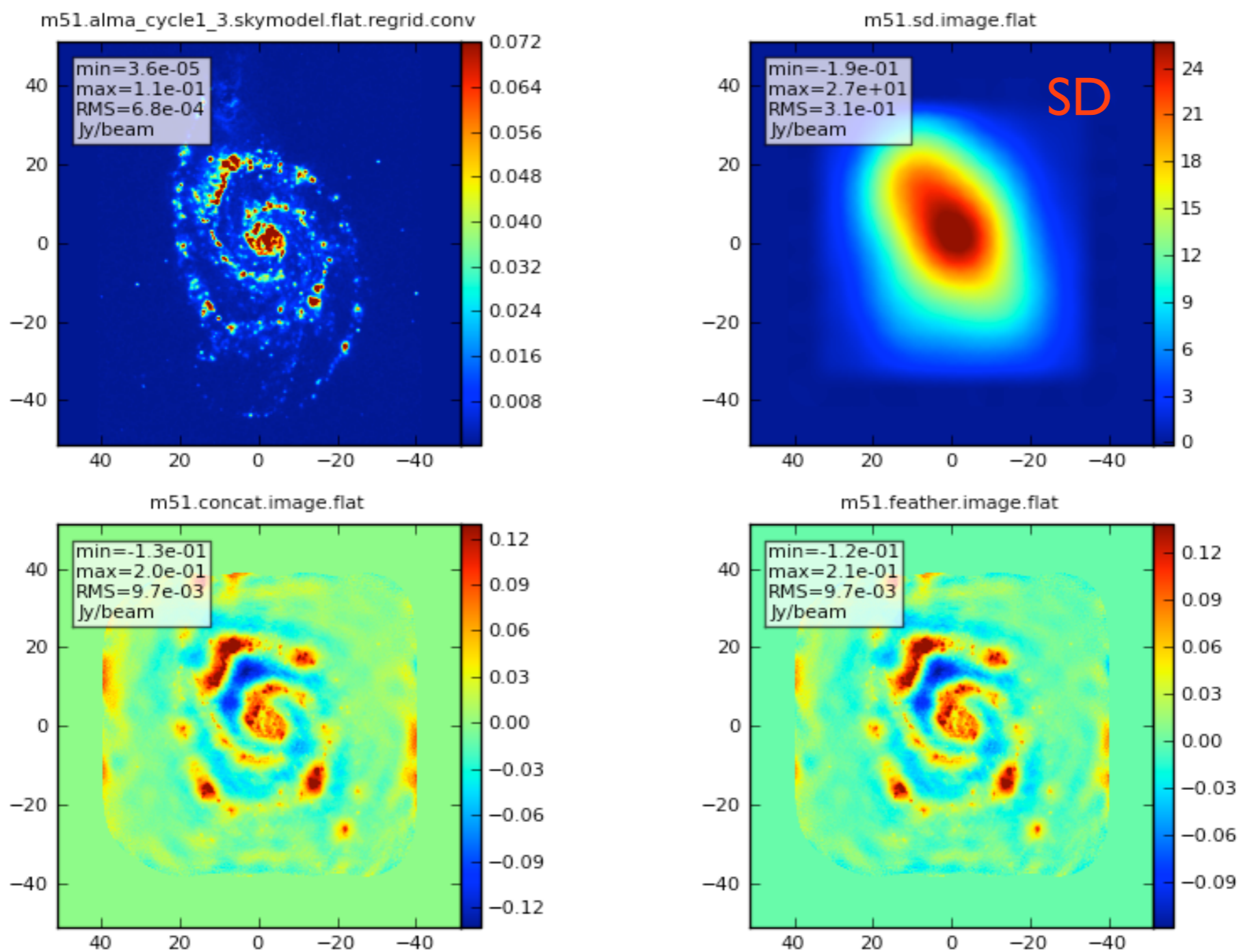
input model and pointing



M5 I input

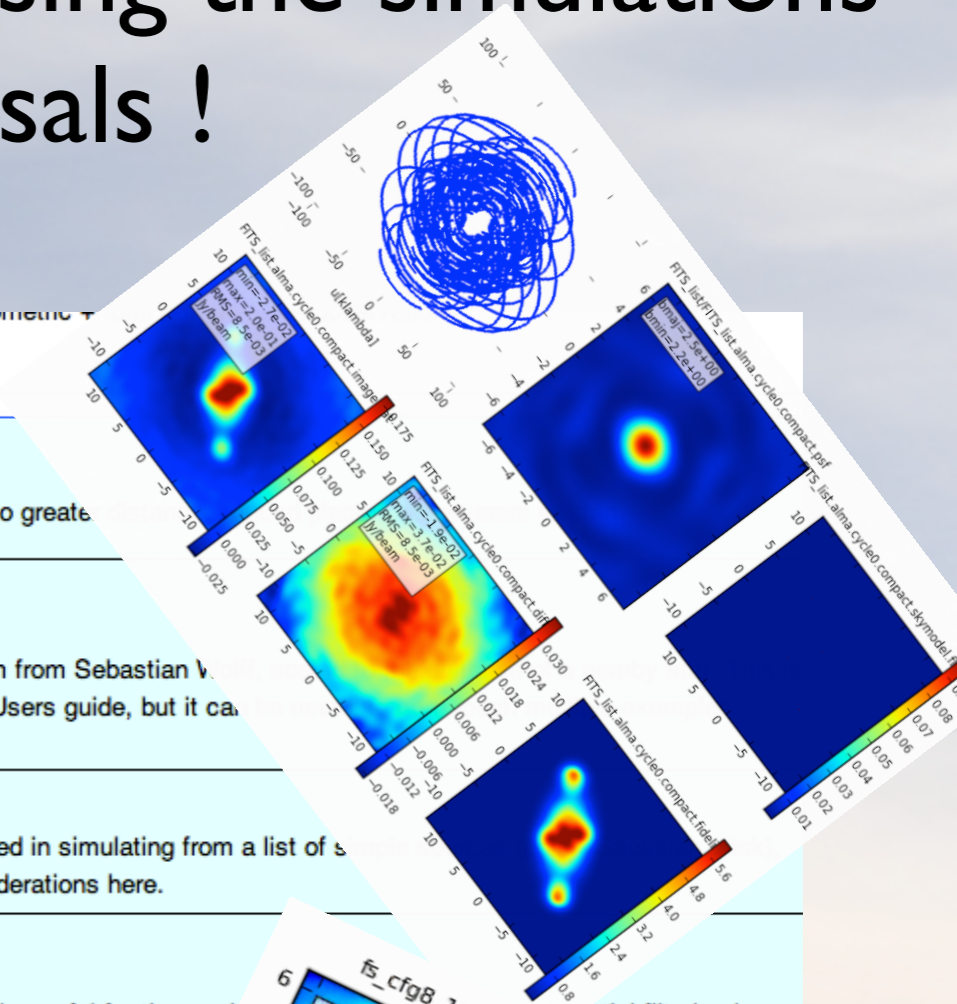
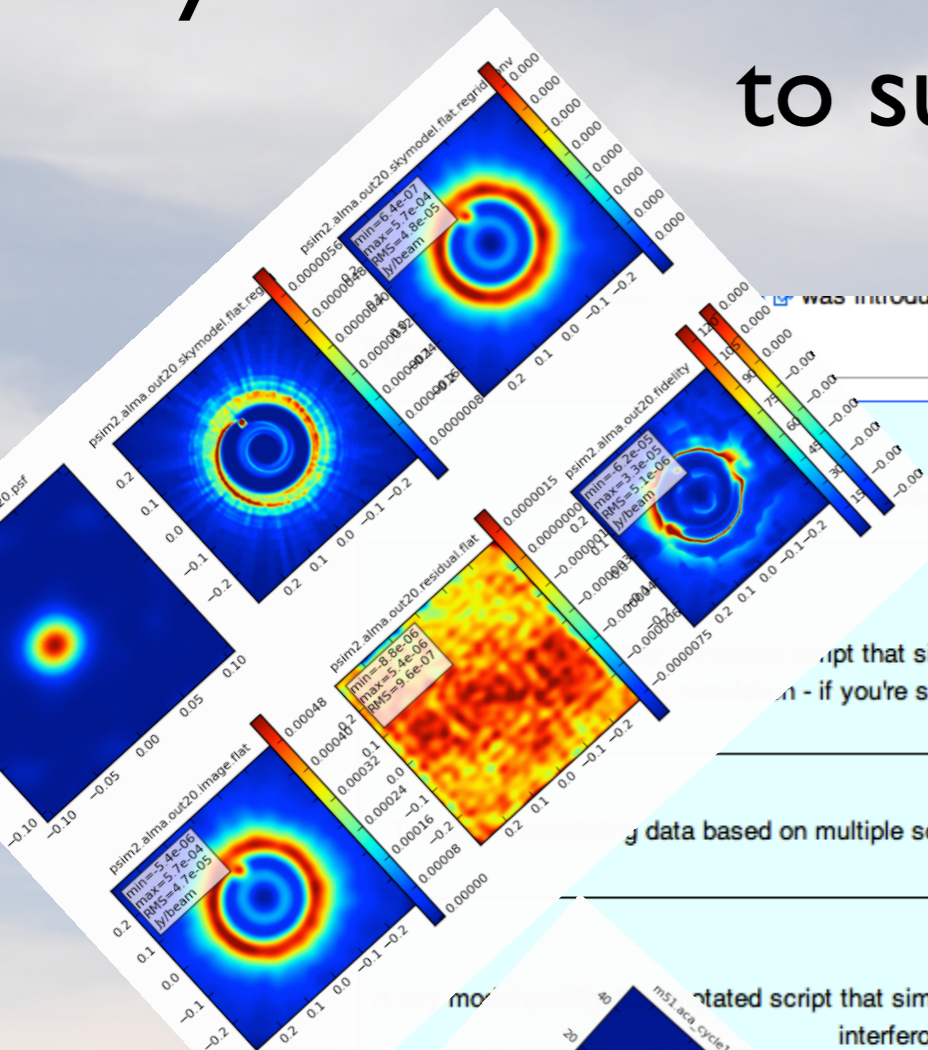


Simalma - combination of SD/ACA/ALMA during de-convolution



Useful examples

Try out the simulators and have fun using the simulations to support your proposals !



Simulation Guide for New Users (CASA 4.3)

a Spitzer SAGE 8 micron continuum image of 30 Doradus and scales it to create

Protoplanetary Disk Simulation (CASA 4.3)

script that simulates a protoplanetary disk. Uses a theoretical model of dust continuum from Sebastian V. - if you're short on time, you probably don't need to go through this one and the New Users guide, but it ca

Simulation Guide Component Lists (CASA 4.3)

g data based on multiple sources (using both a FITS image and a component list). If you are interested in simulating from a list of s rather than or in addition to a sky model image, then read the considerations here.

Einstein-Face (CASA 4.3)

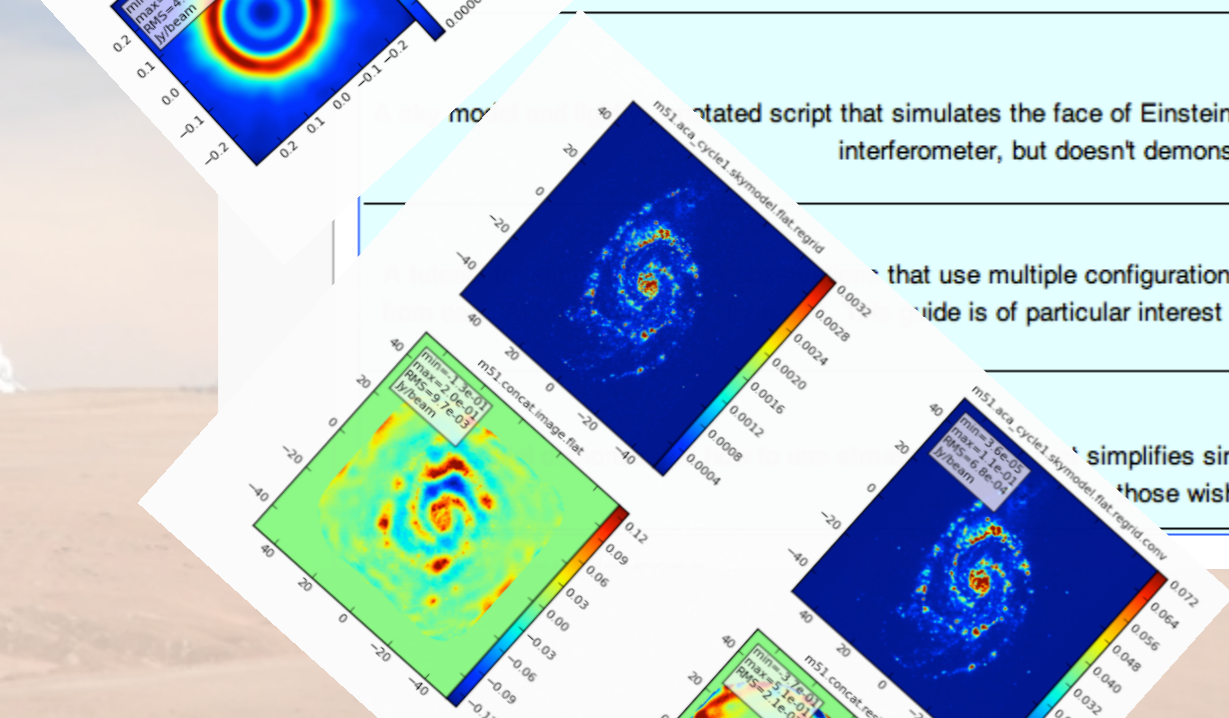
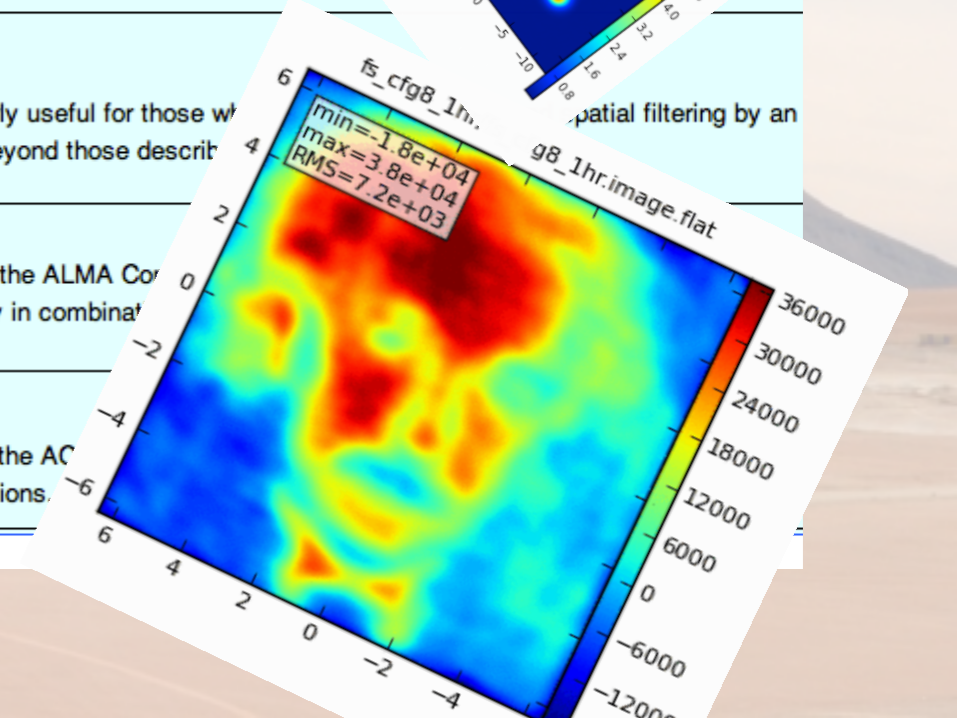
mo- nated script that simulates the face of Einstein as seen by ALMA. This simulation is particularly useful for those w/ nterferometer, but doesn't demonstrate new capabilities of the simulation tasks beyond those descri

ACA Simulation (CASA 4.3)

that use multiple configurations or use the 12-meter array in combination with the ALMA Cor ide is of particular interest to those wishing to explore using the 12-m array in combina from multiple 12-m array configurations.

Simalma (CASA 4.3)

simplifies simulations that include the main 12-m array plus the AC those wishing to explore multi-component ALMA observations.



more complicated sims

Use the simulator tool for:

phase delay variations

gain fluctuations and drift

cross-polarization

(soon also bandpass and pointing errors)

more flexibility in adding thermal noise

Thermal noise addition

<https://safe.nrao.edu/wiki/pub/ALMA/SimulatorCookbook/corruptguide.pdf>

simple: specify **simplenoise="1Jy"** to get random Gaussian noise with 1Jy RMS

- **tsys-atm:** use environment temperatures, antenna parameters, and the aatm library to create a model of the troposphere and add random noise of the appropriate magnitude to the visibilities
- **tsys-manual:** specify atmospheric brightness temperature and optical depth yourself (rather than let aatm calculate it for you) and apply noise of the corresponding magnitude.

