

## ALMA Polarization Commissioning over next Year

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- Some early commisioning results
  - Fixed errors in XY YX connections
    - Eg labeled X3 with Y4 and X4 with X3 was really X5 with Y2 and X4 with Y6.
    - Nothing made sense.
  - Feed orientation of all antennas at any band the same
  - Place feed orientation of feeds in data base.

## Recent Commissioning Activities

- Nagai showed examples of D-term versus time, Frequency and primary beam location.
- George over last two days showed robustness of the Casa software and other procedures to  $<0.3\%$  level
- -- Your successful running of scripts shows that it is not 'too difficult' to obtain accurate I,Q,U,V images.
- ALMA has no remaining concerns about offering Continuum pol at standard freq in bands 3,6,7, Single pointing with 65% Power sensitivity.
- Need one or two Science Verification Observations and The associated casa guide.  
Any suggestions? 3C286 one of them?
- OT implementation has been defined for cycle-2

## Suggested and probable activities over next 6-12 months.

- Confirm observing mode of 'normal' phase referencing
  - Plus ~6 observations of strongly polarized sources
  - Every 30 min.
    - a. Enough parallactic angle coverage? Source dec Range of -60 to +10 okay.
    - b. How to find/monitor ~15 quasars with at least 4% Polarization, greater than ~0.5 Jy.
- Baselines out to 3 km. How to deal with weak phase calibrators with unknown polarization, exp band 7
- Pushing angular size limit to ~40% of power pattern. Does Mosaicing help.
- Always observe in all 4-Stokes unless the spectral Resolution is needed.

## Suggested and Possible activities for cycle-3.

- Begin testing more efficient pol observations
  - Egs. Short obs of polarized cal and unpolarize cal
- Monitor D-term solutions and check long-term stability
  - Can we stop determining D-terms for standard freqs?
- -- Analysis of high spectral resolution D-term determination.
  - Can we obtain TDM results and interpolate to FDM?
- Deeper mosaic studies:
  - ACA + ALMA mosaics should not be a major problem.
  - Add in SD polarization. Will be challenging.
- Circular polarization accuracy and observing/reduction
  - Methods needed.